



# **Application of Novel Hybrid Mesh Generation Methodologies for Improved Unstructured CFD Simulations**

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Deutsches Zentrum  
für Luft- und Raumfahrt e.V.  
in der Helmholtz-Gemeinschaft

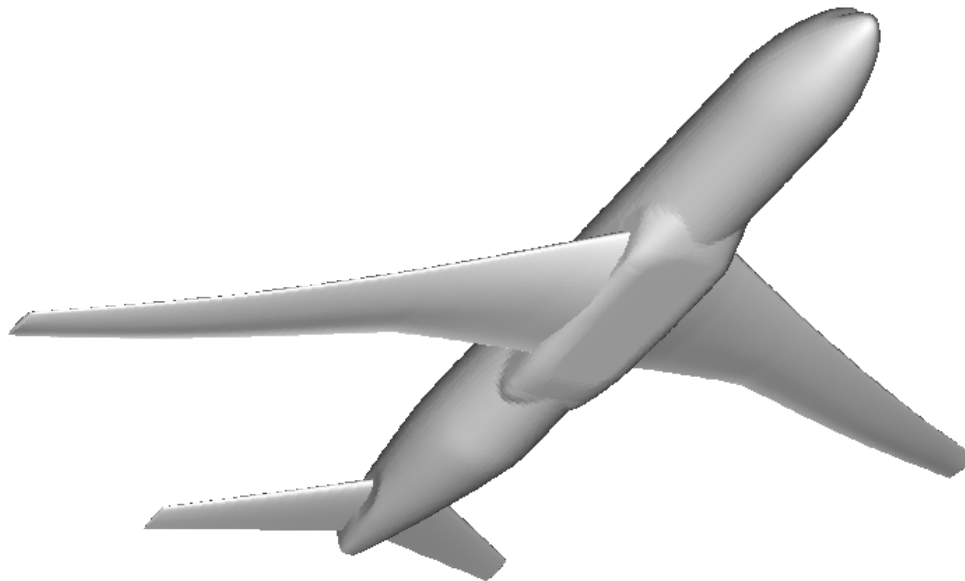
# Overview

- Introduction & Cases
- Computational Methods
- Initial Results
- Up-Front Grid Improvements
- Grid Family
- Adjoint-based Grid Improvements
- Workshop Summary
- Post-Workshop Activities
- Conclusions



# Introduction

- NASA common research model; cruise configuration of conventional civil transport aircraft
- WBT-configuration selected as experimental basis for the fourth Drag Prediction Workshop (DPW4)
- “Natural application” of Solar grid generation package due to available philosophy files, which automate the definition of sources



# Cases

- Workshop held in San Antonio, TX, on June 20-21 2009
- **case 1** – consisting of two sub-cases
  - 1) grid convergence study at  $M = 0.85$ ,  $C_{\text{lift}} = 0.5$ ,  $Re = 5 \cdot 10^6$ 
    - tail incidence angle  $i_H = 0^\circ$
    - coarse **medium** and fine grids
  - 2) downwash study at  $M = 0.85$ 
    - drag polars for  $\alpha = 0.0^\circ, 1.0^\circ, 1.5^\circ, 2.0^\circ, 2.5^\circ, 3.0^\circ, 4.0^\circ$
    - tail incidence angles  $i_H = -2^\circ, 0^\circ, +2^\circ$ , and tail off
    - medium grid
    - trimmed drag polar derived from polars at  $i_H = -2^\circ, 0^\circ, +2^\circ$
    - delta drag polar of tail-off vs. tail-on
  - three turbulence models (SAO, SST, RSM)
- case 3 – Reynolds number effects (WT to flight scaling)
  - Same configuration & settings as case 1.1, but with  $Re = 20 \cdot 10^6$

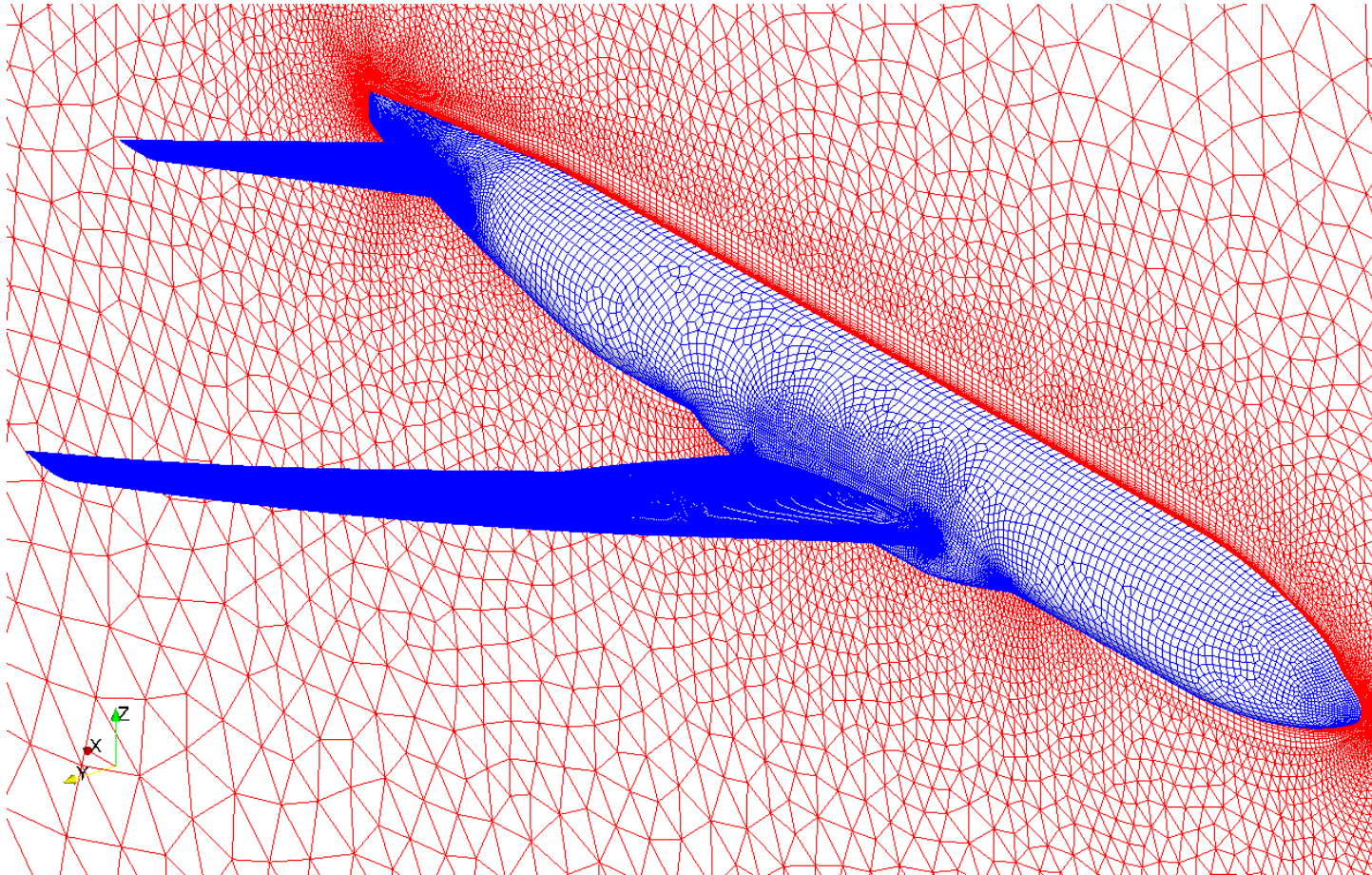


# Computational Methods

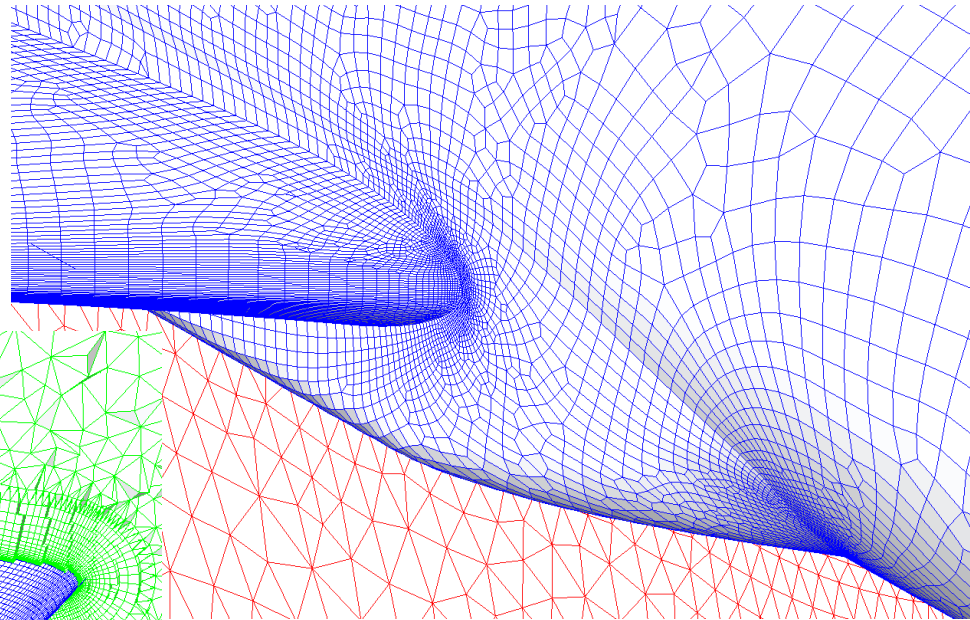
- DLR **TAU** code (CFD) and ARA/BAE-ATC **Solar** (grid gen.)
- Compressible RANS
- Unstructured, finite volume, node-centered, (Full-NS)
- Central scheme with Jameson-type scalar dissipation;  $k_2=1/4$   
 $k_4=1/64$
- Backward-Euler, LUSGS
- 4w multigrid / sg
- SA(RC), Menter  $k-\omega$  SST, SSG/LLR- $\omega$  RSM
- Unstructured, quad/hexa-dominant
- Advancing-layer/front
- Octree-based background grid for steering Delaunay tetra generation
- Anisotropic surface quads
- Adaptive, variable expansion ratio
- No user-specified amount of wall-normal layers

# Initial Results

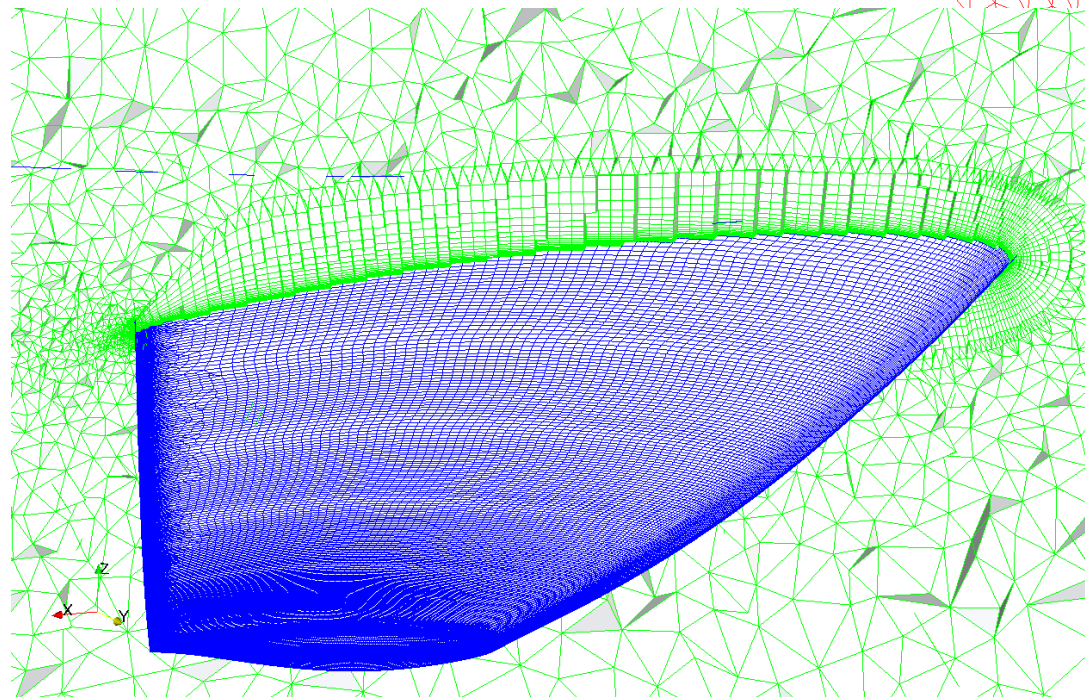
- Initial medium grid (~10 mio. Points) generated in ½ day, using default philosophy files



# Initial Results



Wing-fuselage fairing

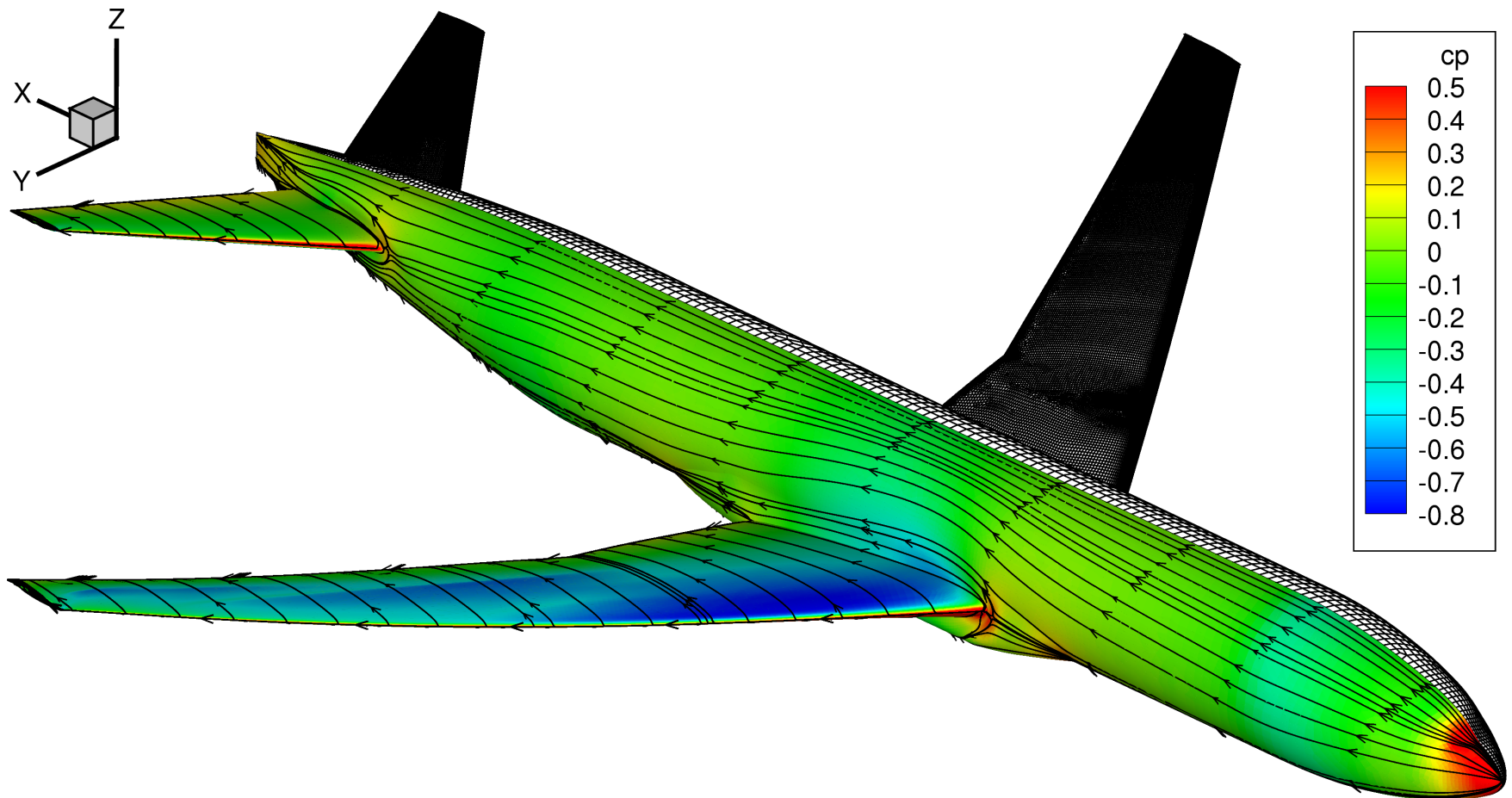


Field-cut around outboard wing



# Initial Results

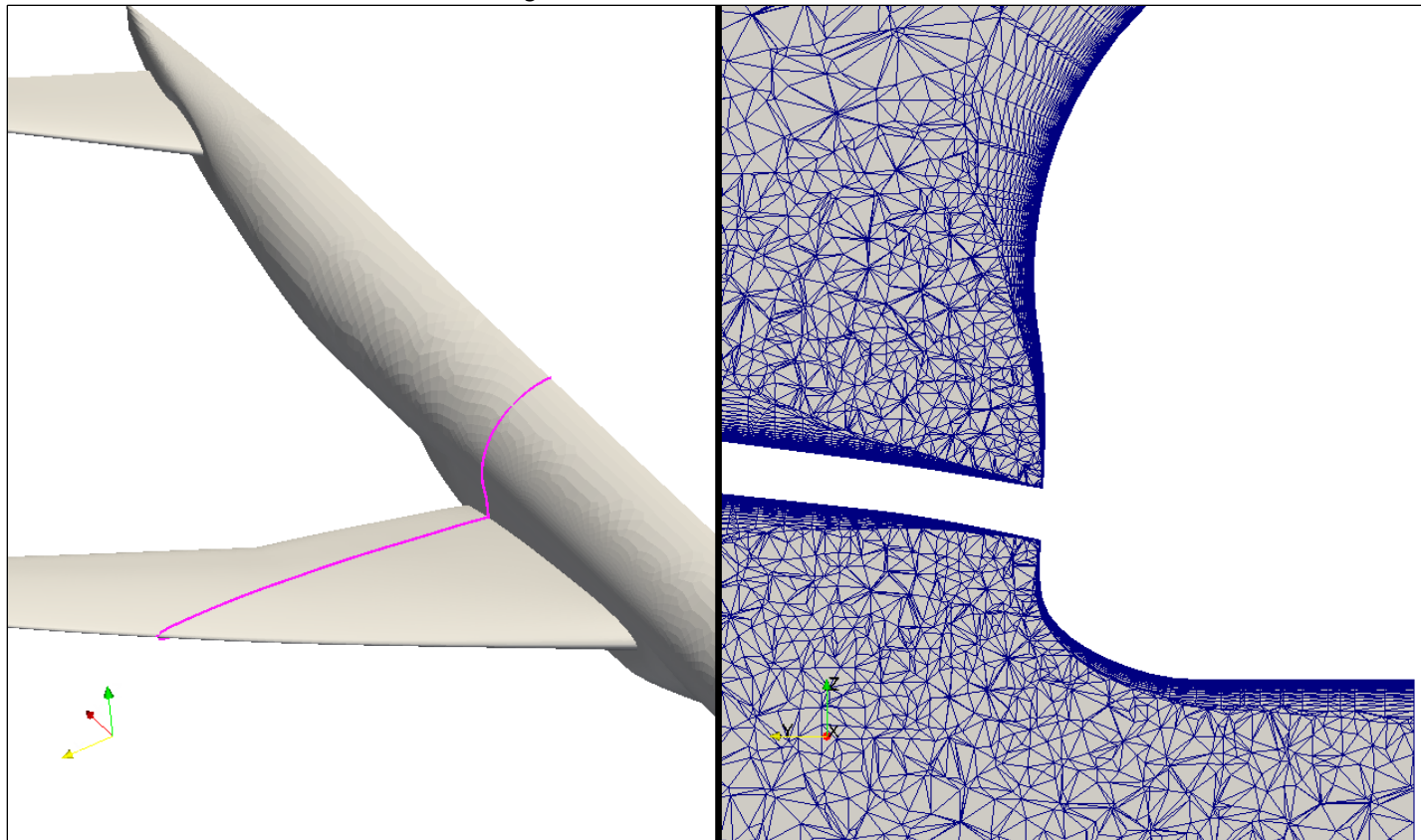
➤ Overnight initial results; acceptable solver convergence





# Up-Front Grid Improvements

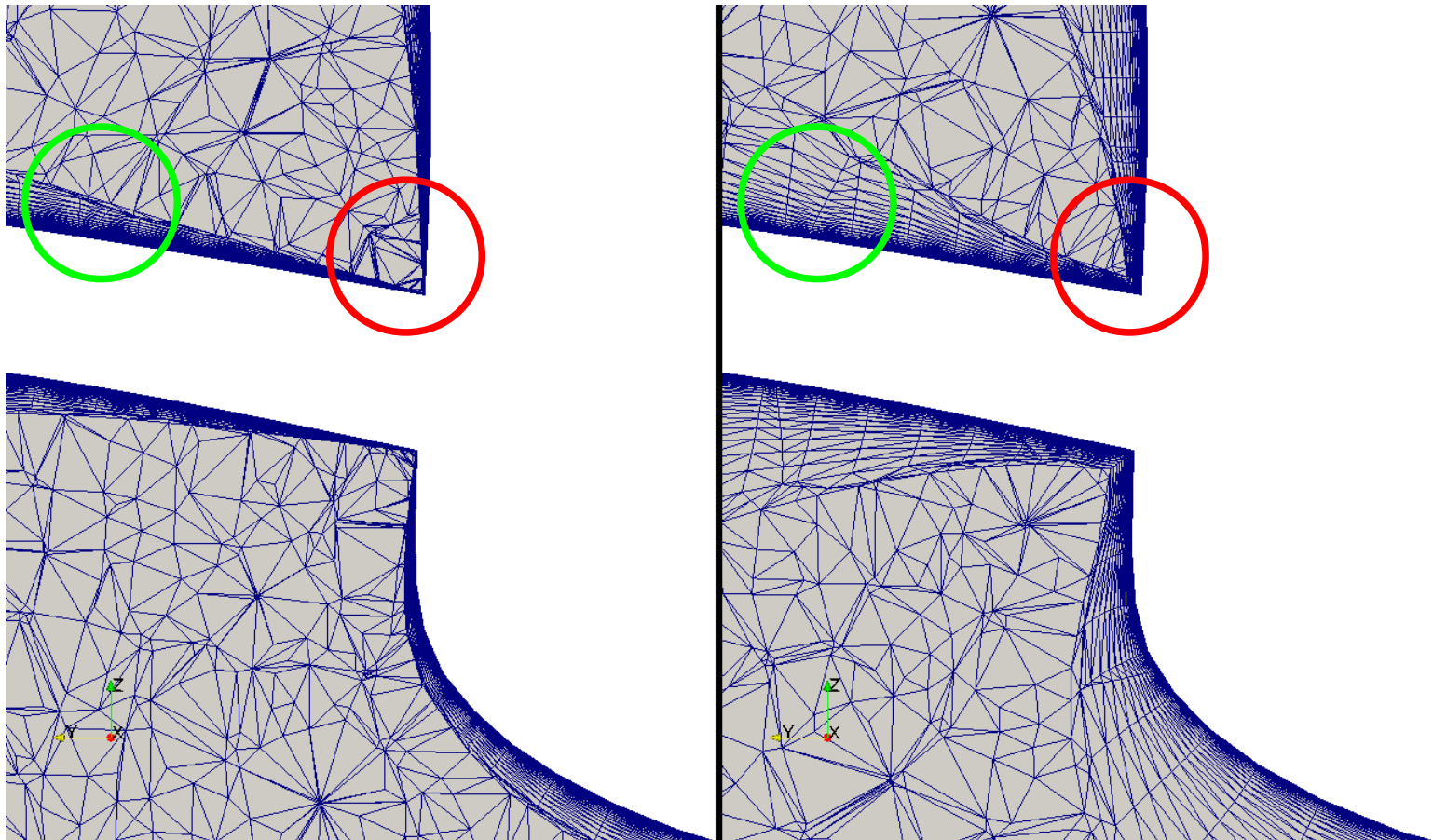
- Strict application of DPW committee gridding guidelines leads to contracted near-field layer at concave corners



Cut at  $x=1400''$

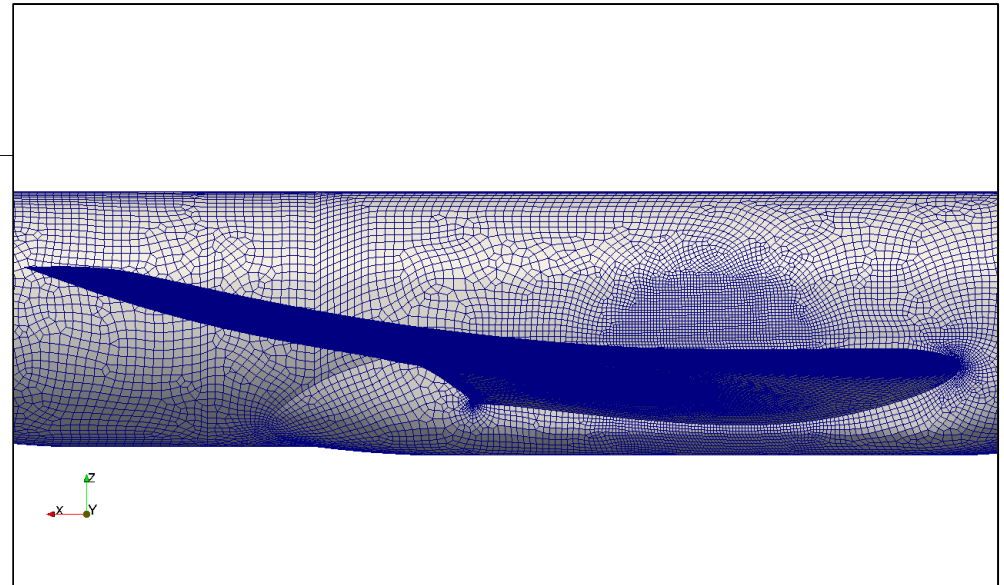
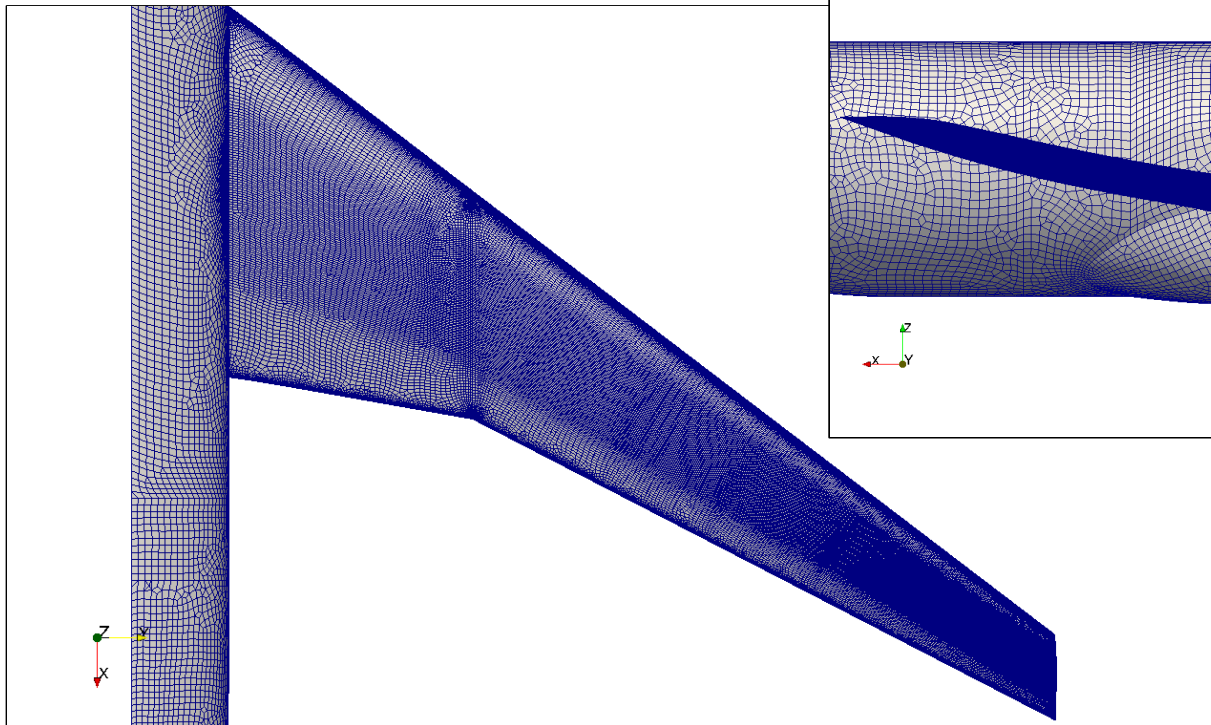
# Up-Front Grid Improvements

- Increase of surface grid size at junctions mitigates, but does not resolve the problem



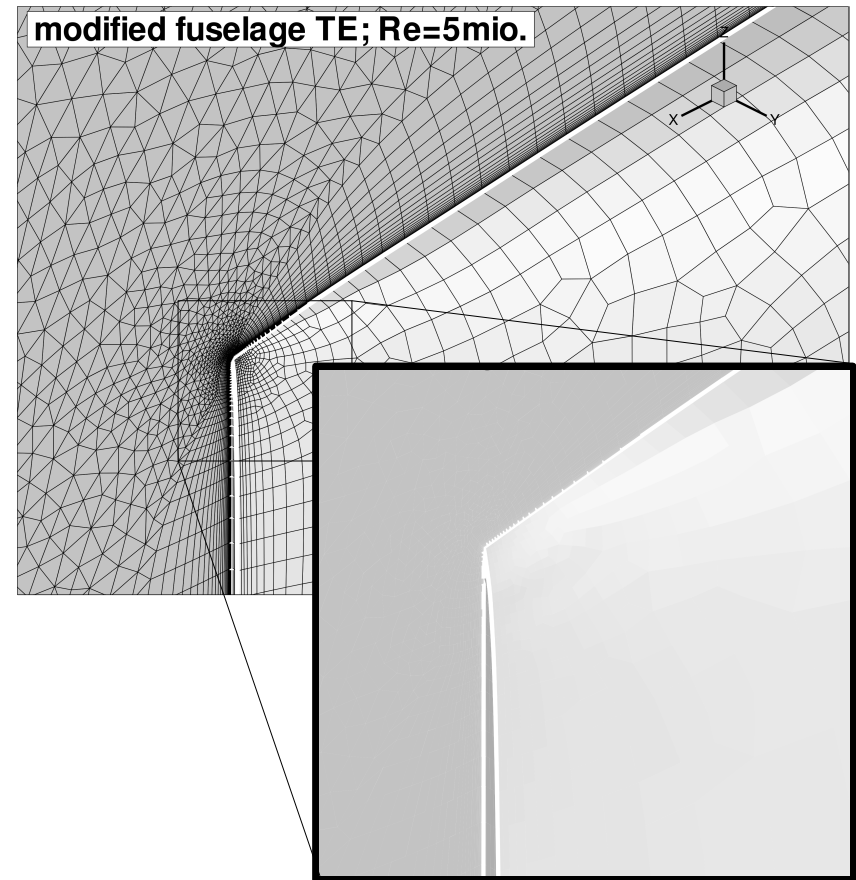
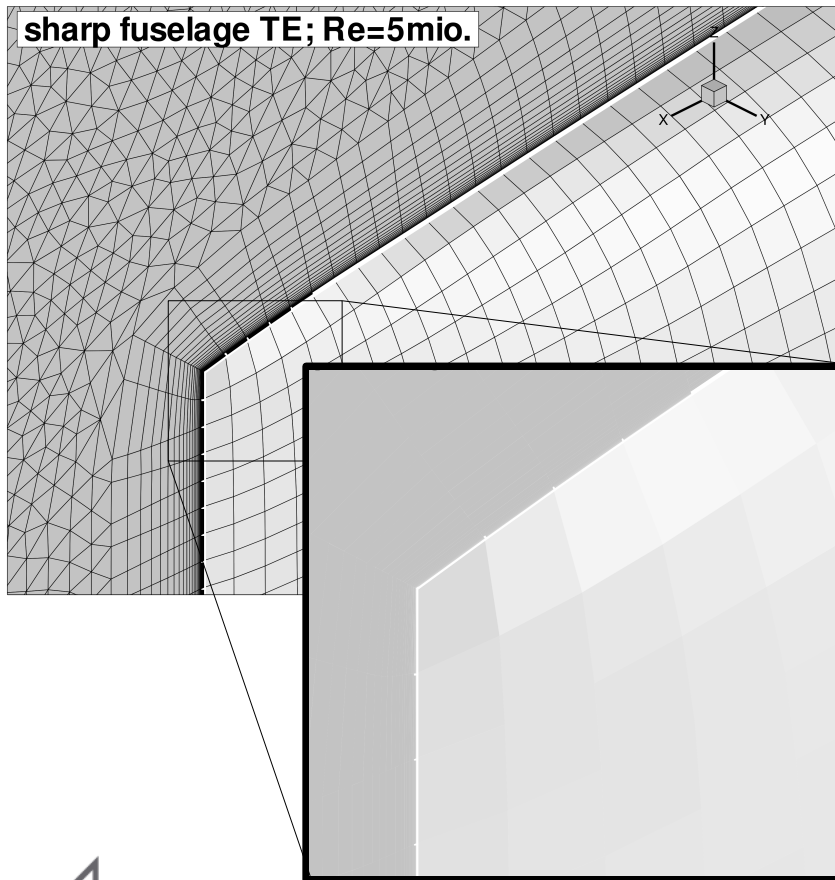
# Up-Front Grid Improvements

- After initial runs with medium grid, additional anisotropic source added at mid-chord, over entire span, to better resolve shock region



# Up-Front Grid Improvements

- Modification of fuselage trailing edge
  - Sharp → blunt; through cut at  $x=2561.5$  inch (bTE: 0.5 inch)





# Grid Family

- Starting from medium source distribution, generation of one level of finer and coarser grids
  - Source sizes scaled by a factor of  $\sqrt[3]{3}$  ( $\approx 1.4422$ ), affecting both surface and volume meshing
  - Influence radii ( $r_1$  &  $r_2$ ) not changed, being coupled to geometry
  - Consistent scaling of expansion ratio, to keep the total near-field layer extent similar between grid levels

$$\text{Total layer thickness} = \sum_{i=0}^n a \cdot q^i = a \cdot \frac{1 - q^{n+1}}{1 - q}$$

$a$ : first layer spacing  
 $q$ : expansion ratio  
 $N$ : Number of layers ( $n+1$ )

Grid level 2 being finer than grid level 1:  $a_2 = a_1 / \sqrt[3]{3}$  ,  $N_2 = N_1 \cdot \sqrt[3]{3}$

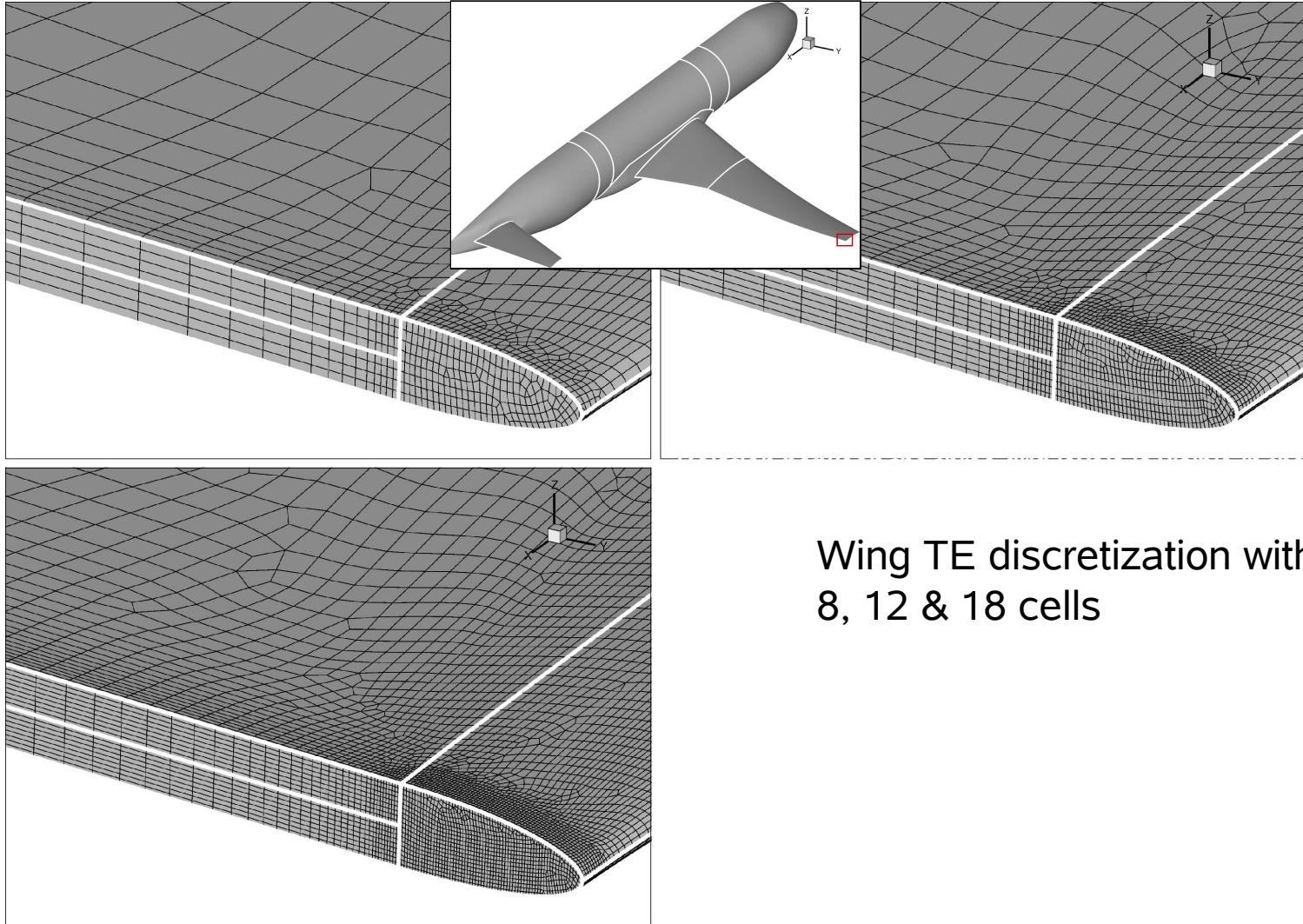
$$a_1 \cdot \frac{1 - q_1^{N_1}}{1 - q_1} = a_2 \cdot \frac{1 - q_2^{N_2}}{1 - q_2}$$

# Grid Family

## ➤ Final grid family

Grid	Coarse	Medium	Fine
# cells on blunt TE	8	12	18
Surface point	$130 \cdot 10^3$	$271 \cdot 10^3$	$566 \cdot 10^3$
Max. # of wall-normal layers	30	42	60
Expansion ratio	1.3009	1.2	1.135
# points in nearfield	$3.47 \cdot 10^6$	$9.94 \cdot 10^6$	$28.69 \cdot 10^6$
# of tetra cells	$5.31 \cdot 10^6$	$14.31 \cdot 10^6$	$38.58 \cdot 10^6$
<b>Total # of points</b>	<b><math>4.07 \cdot 10^6</math></b>	<b><math>11.70 \cdot 10^6</math></b>	<b><math>34.08 \cdot 10^6</math></b>

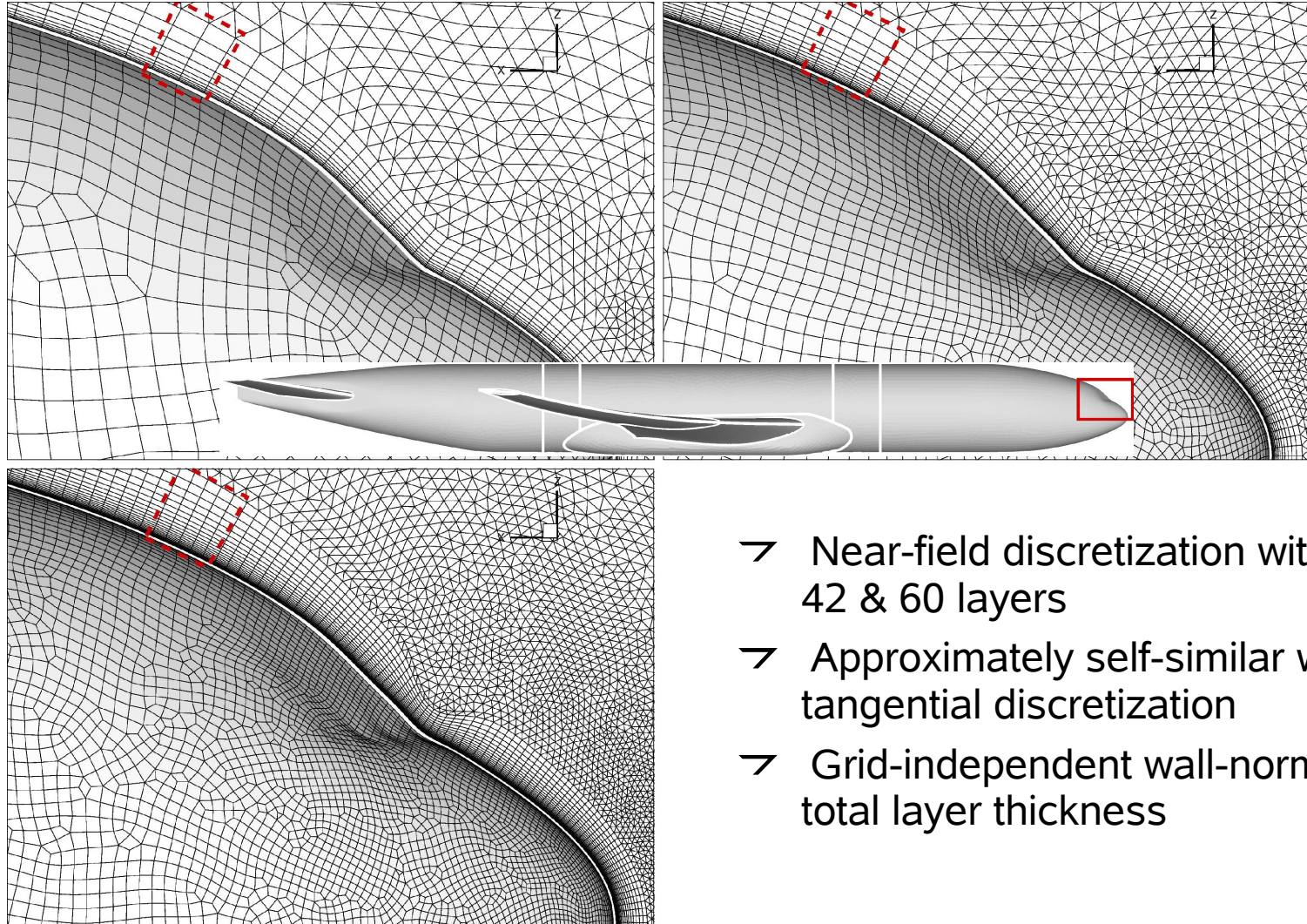
# Grid Family



Wing TE discretization with  
8, 12 & 18 cells



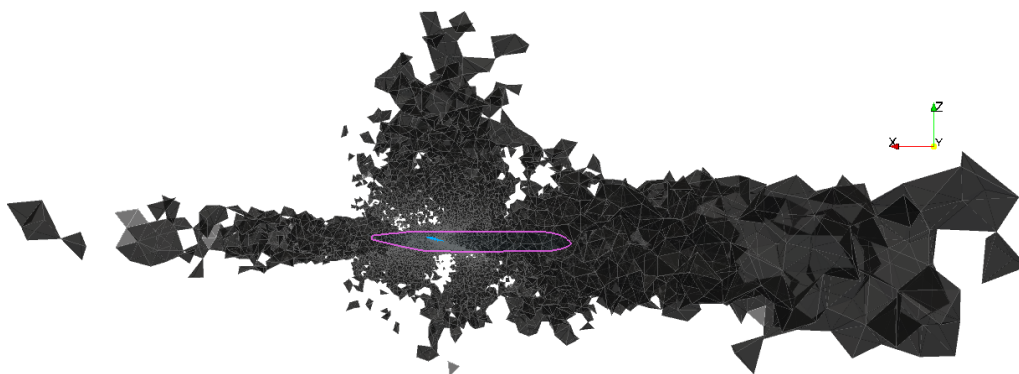
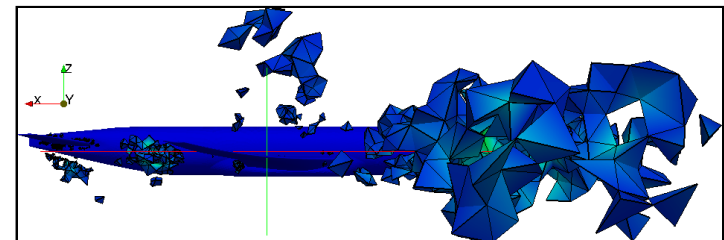
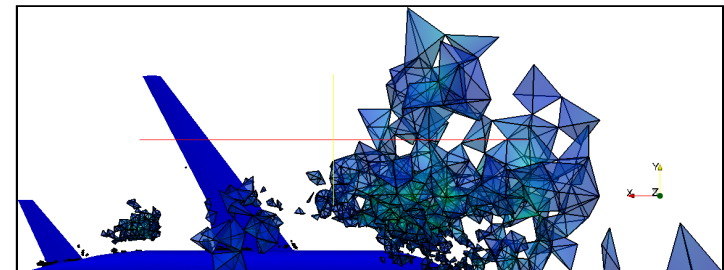
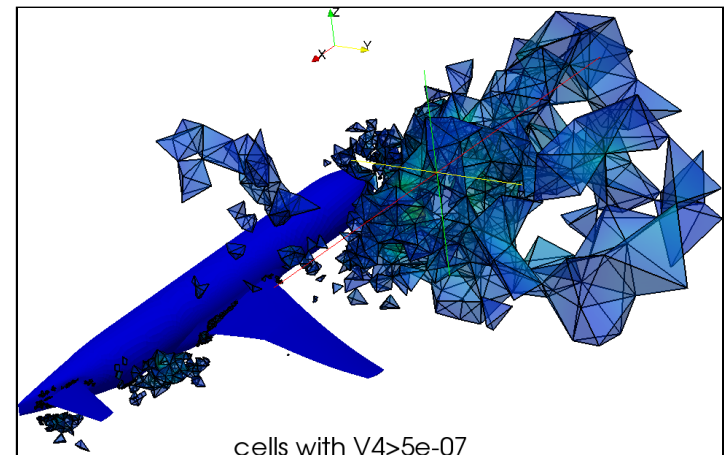
# Grid Family



- Near-field discretization with 30, 42 & 60 layers
- Approximately self-similar wall-tangential discretization
- Grid-independent wall-normal total layer thickness

# Adjoint-based Grid Improvements

- Identification of regions with high influence of artificial dissipation on specific cost function (C-drag/C-my)
  - Variable 4 (V4) used here as sensor from TAU adjoint solver
- Manual, interactive modification and addition of Solar sources

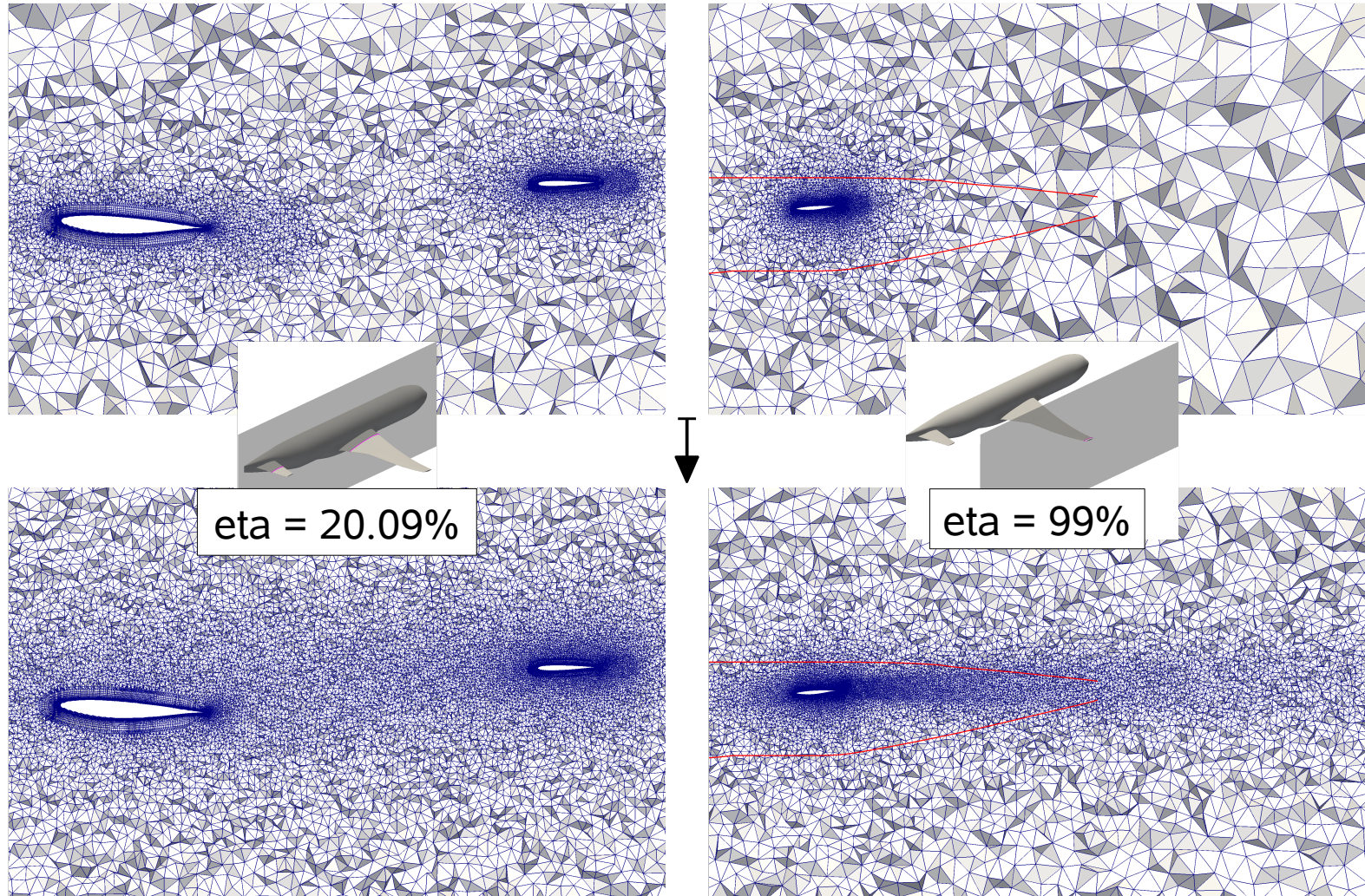


# Adjoint-based Grid Improvements

- Identified problematic regions with remedy
  - Fuselage: spacing x0.5
  - Wing- and HTP-fuselage junction: reduce chord-wise anisotropy, increase spacing x1.5
  - Inner wing wake: spacing x0.5
  - HTP wake: spacing x0.75
  - Upstream of wing: extend upstream wing tri-sources; spacing x8,  $r_1 \times 2$ ,  $r_2 \times 2$
  - Wingtip vortex: add wingtip tri-source

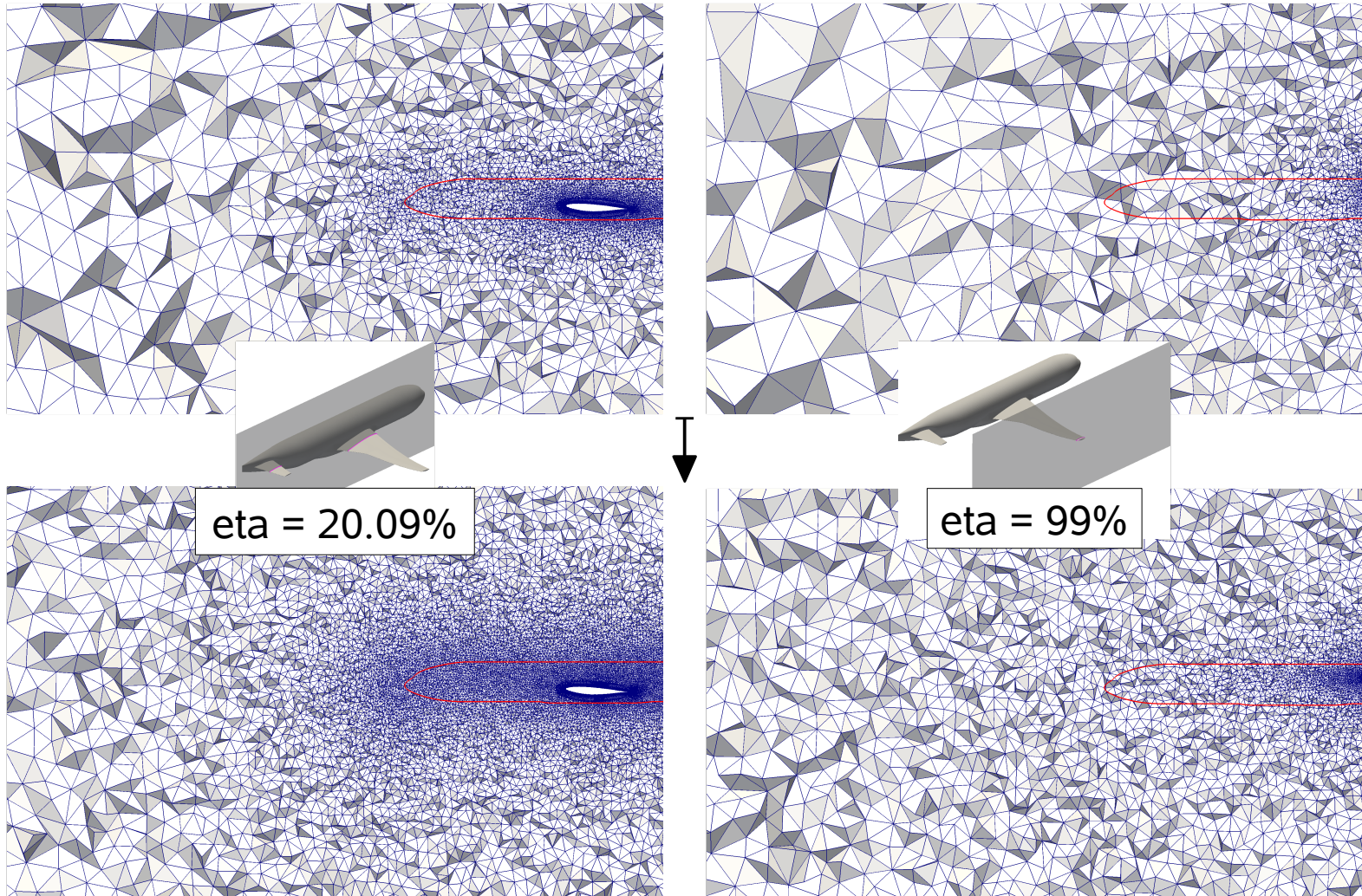


# Adjoint-based Grid Improvements





# Adjoint-based Grid Improvements





# Workshop Summary

- Workshop summary and comparison between all delivered datasets; Solar/TAU: T, C, S (SST, SA, RSM)
- Asymptotic grid convergence
- SST and SA TAU results within scatter of all (SST and SA) structured results

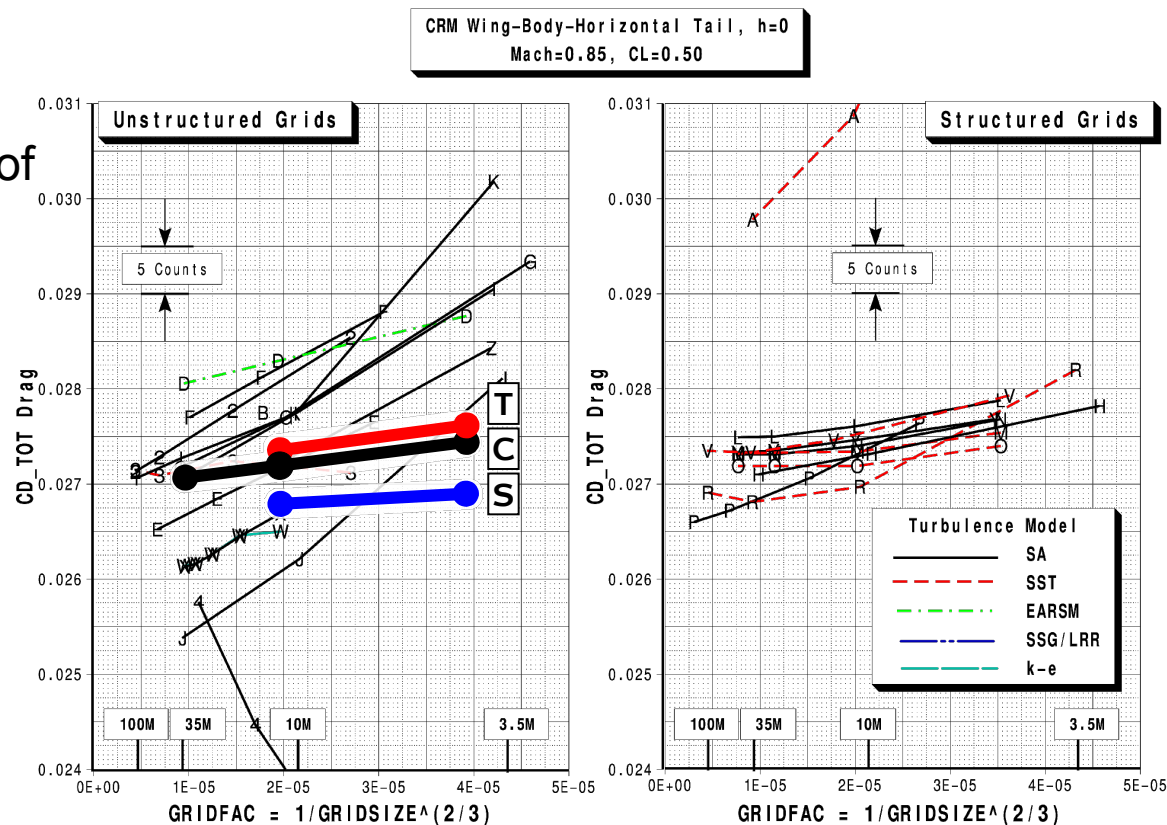


image courtesy: Tinoco et al.

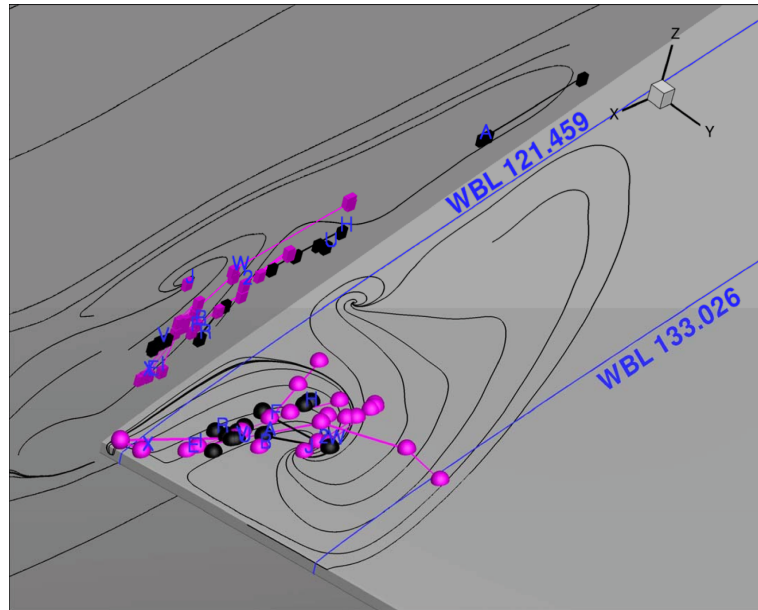
# Post-Workshop Activities

- Wing-body separation bubble at trailing edge
  - Some participants predict one, others don't
  - Solar/TAU (all grids, all turbulence models) no separation

No reported separation

ID      Turb. Model

C	SA
D	EARSM
L	SA
M	SST k-w
N	SA
O	SST k-w
S	SSG/LRR
T	SST



Reported Separation

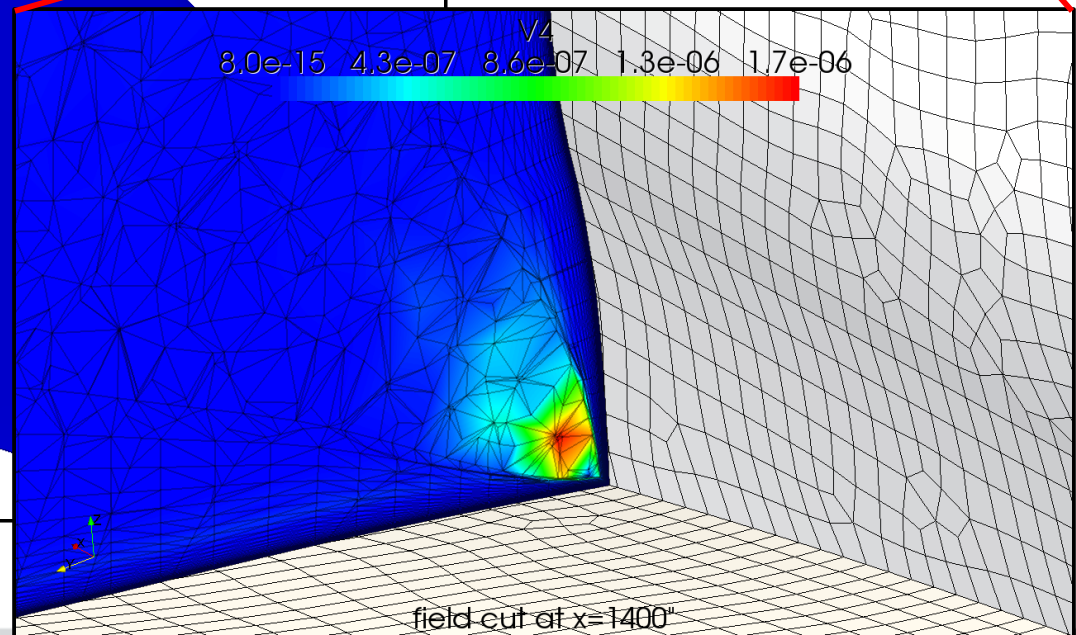
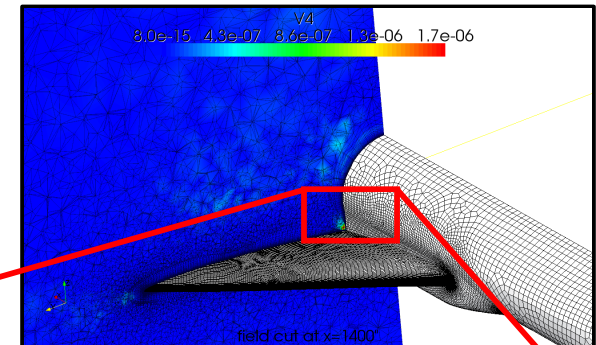
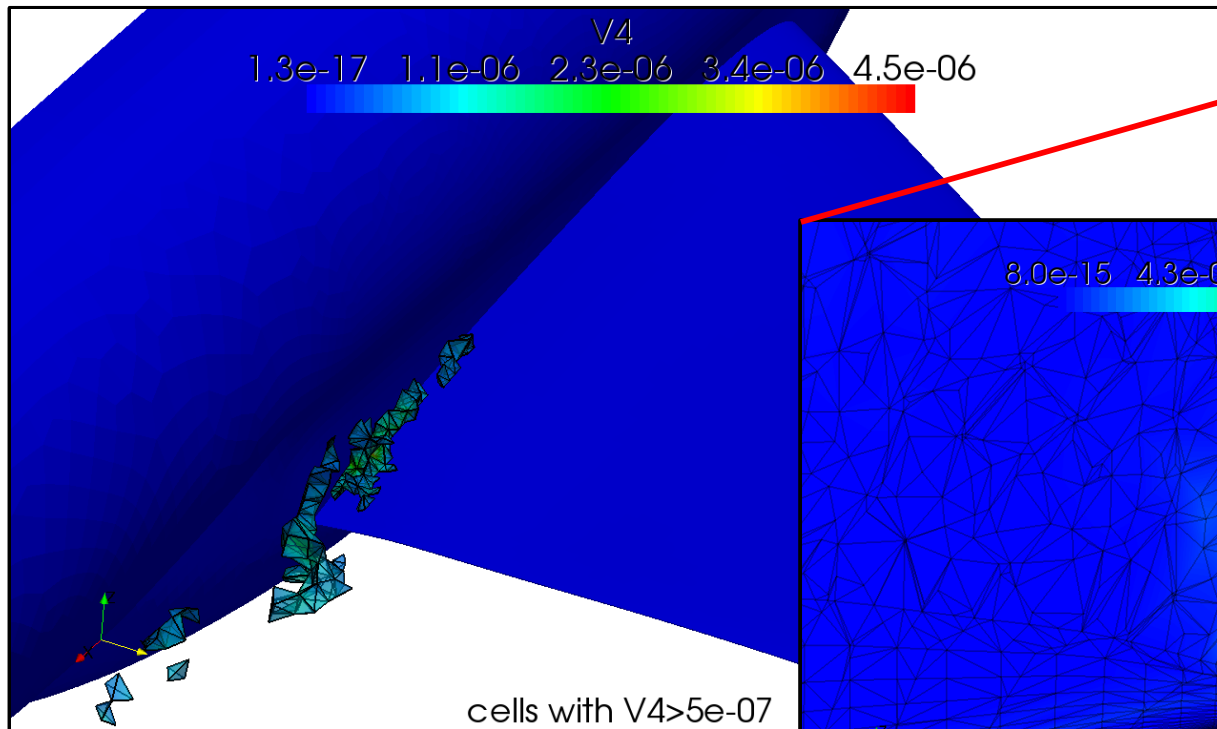
ID      Turb. Model

A	SST k-w
B	SA
E	SA
F	SA
H	SA
I	SA
J	SA
P	SA
R	SST k-w
U	SA
V	SST k-w
W	k-e
X	SA
2	SA
4	SA

image courtesy: Tinoco et al.

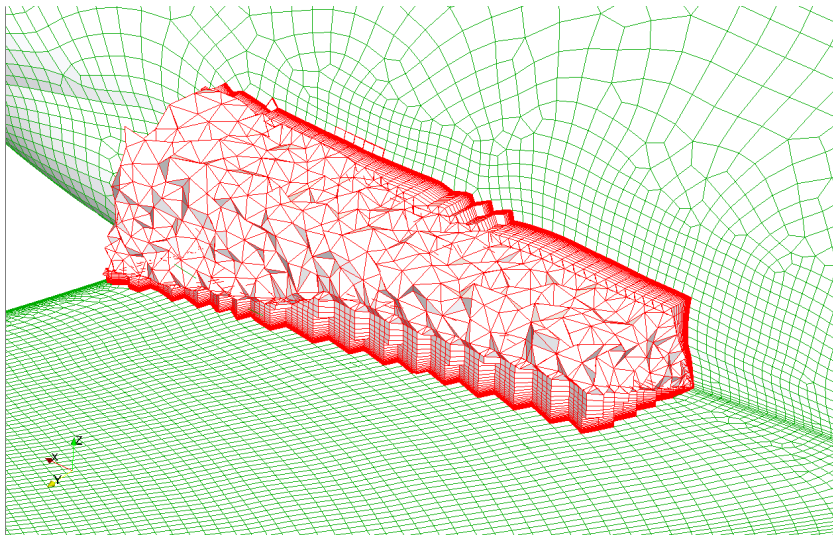
# Post-Workshop Activities

- Hints of insufficient wing-body junction discretization observed from the adjoint solution

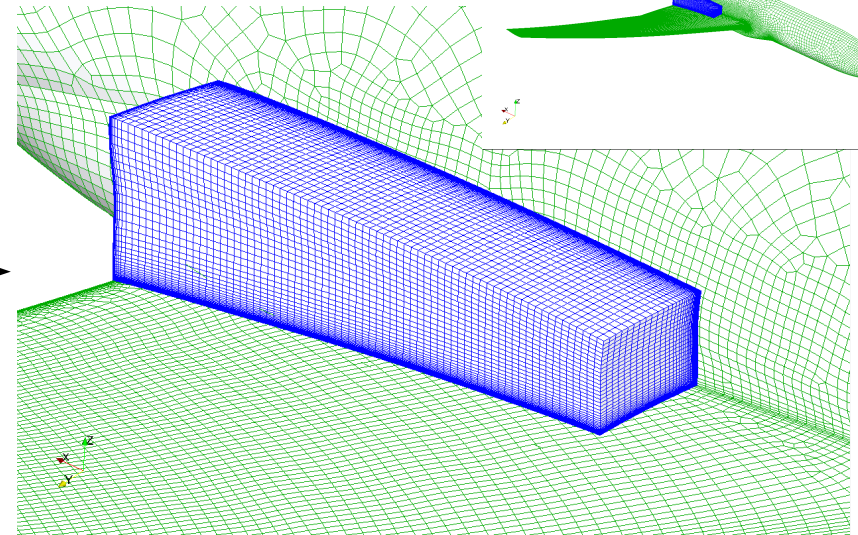
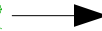


# Post-Workshop Activities

- Solver-side solution approach to junction near-field contraction via overset grid (chimera)
- C-H (C/90°) block initially constrained to wing/body TE region



Cut-out from Solar grid

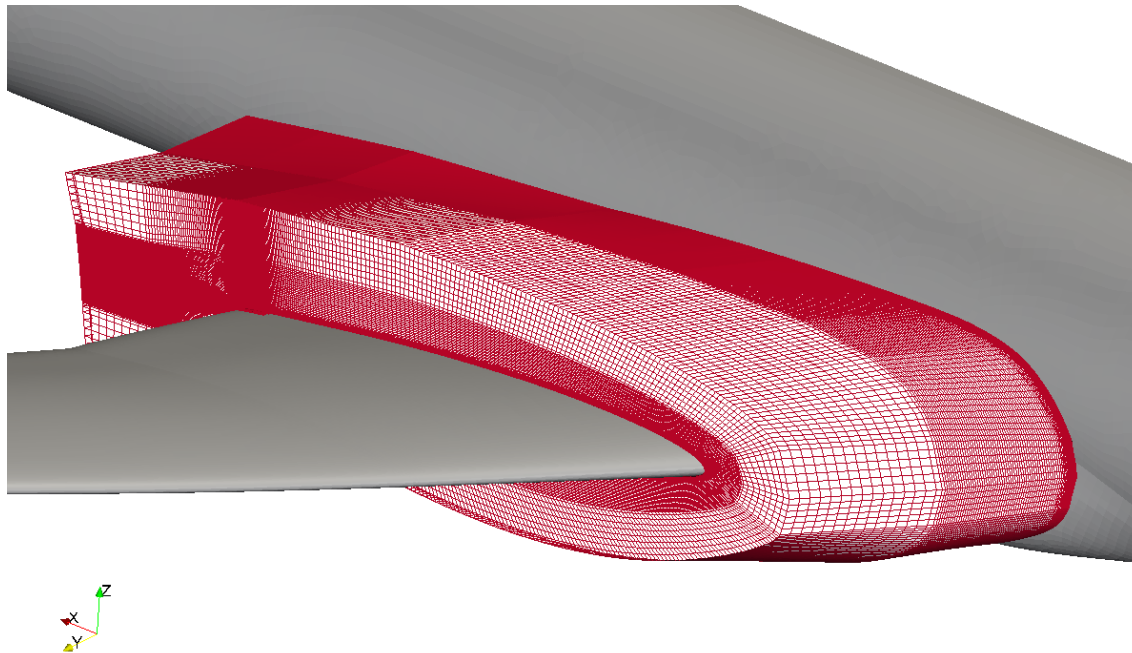


New computational domain



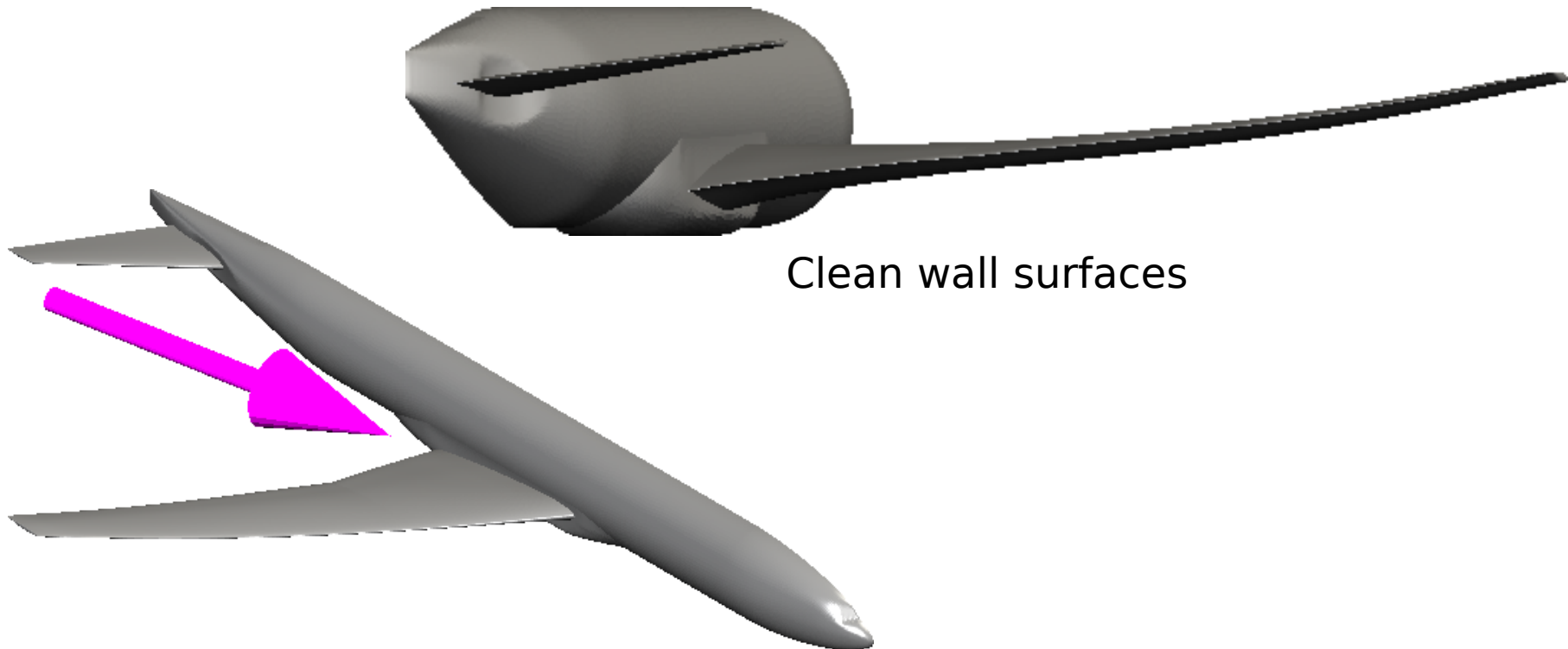
# Post-Workshop Activities

- Based on study of resulting BL profiles from initial hexa block, switch to complete C-H block around entire wing root and wake
  - Same first cell height and expansion ratio at overlap
  - Chimera grid composed of Solar medium grid (11.7 mio. points) + hexa (5.5 mio. Points)



# Post-Workshop Activities

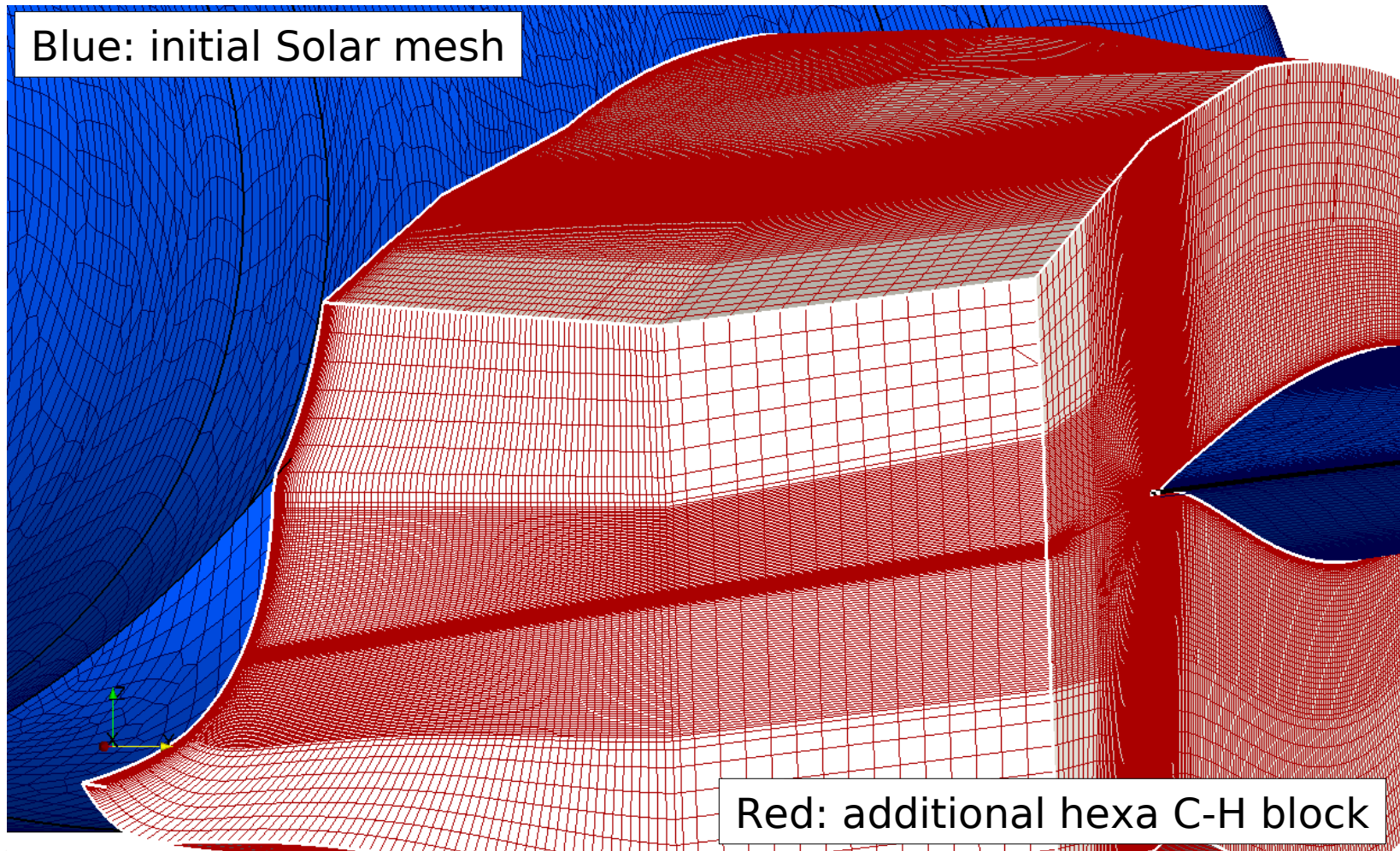
- Orientation for the following slides
  - View from the rear to the front, rotated by  $10^\circ$  around the z-axis



Clean wall surfaces

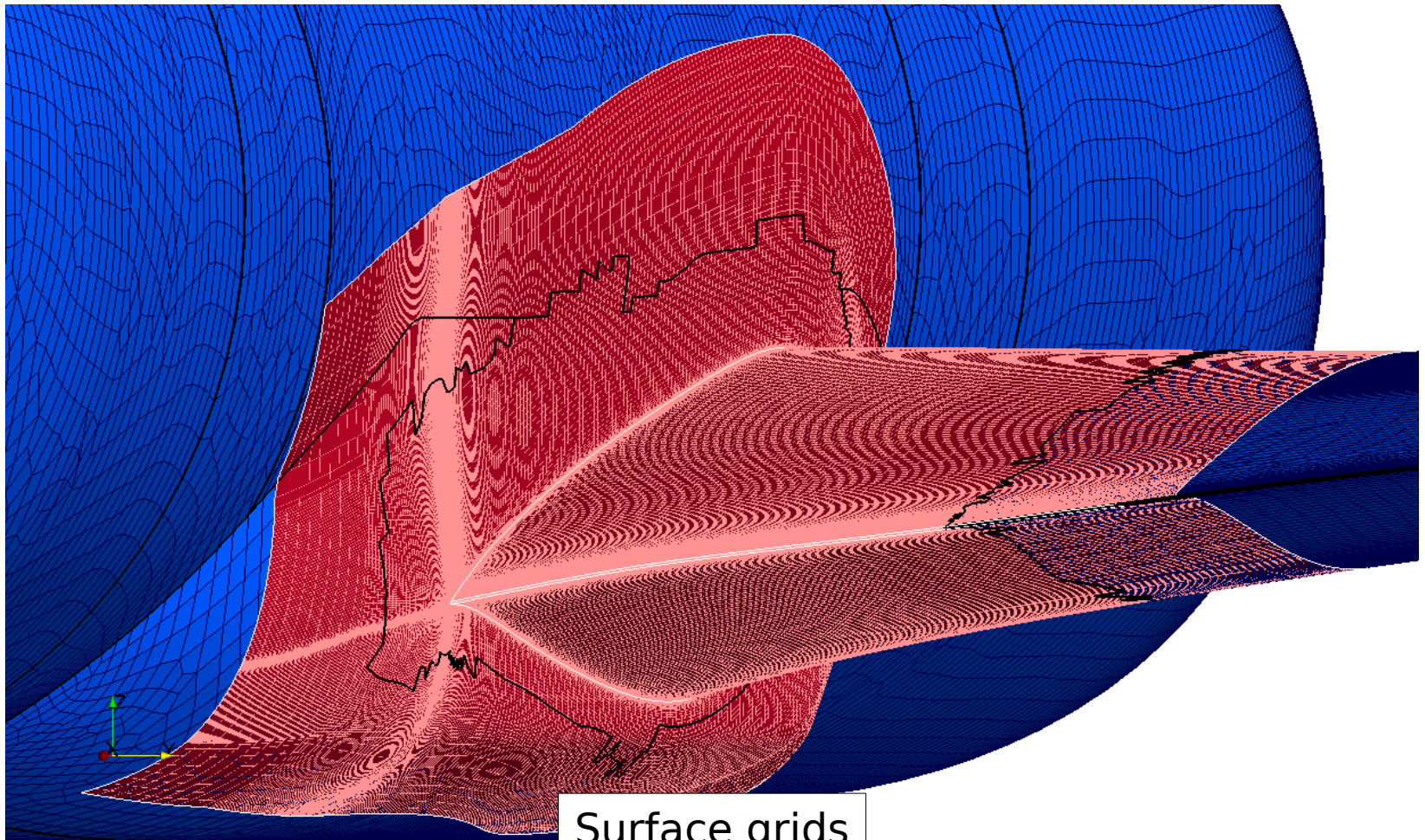
View direction on TE of wing-body junction

# Post-Workshop Activities



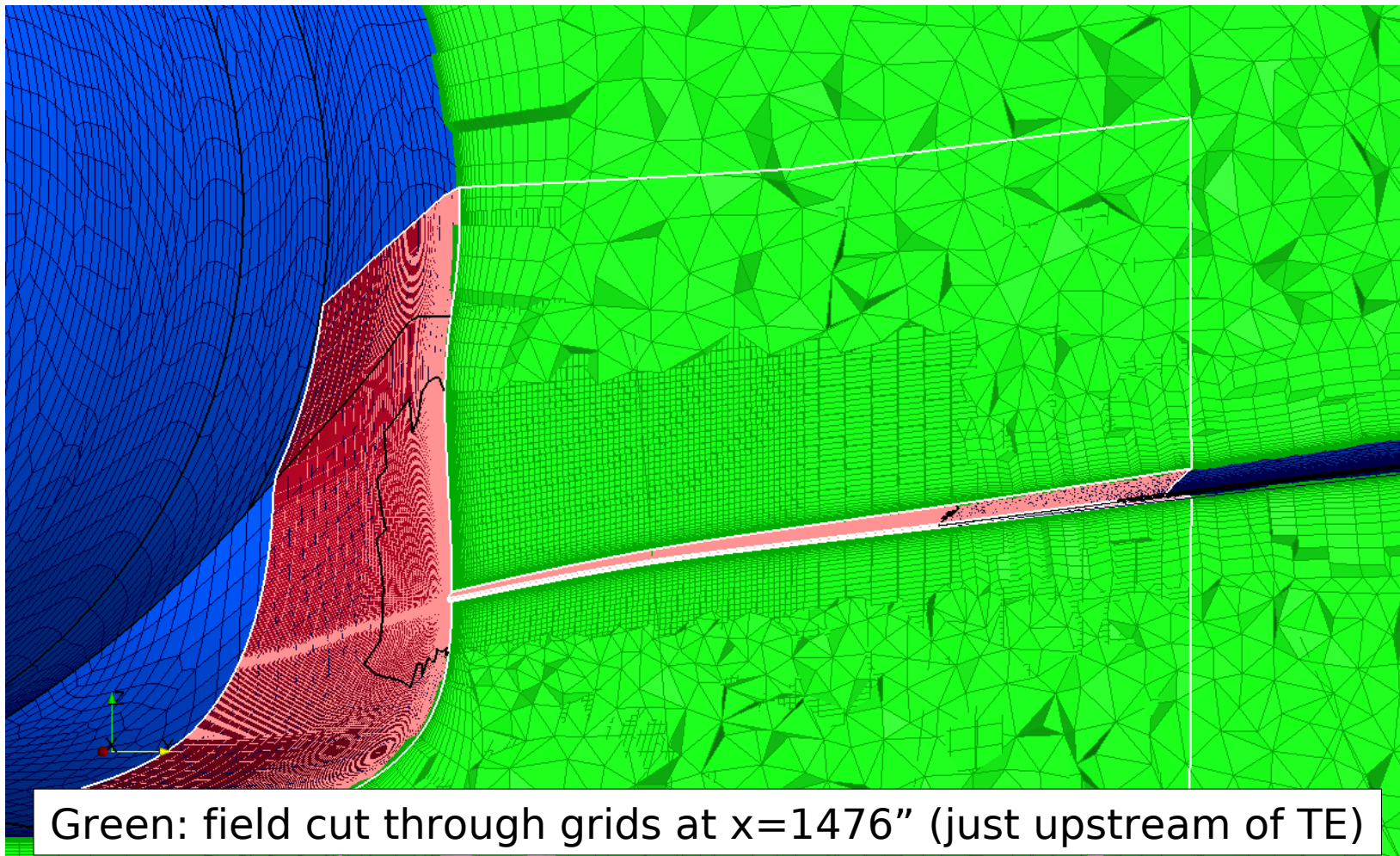


# Post-Workshop Activities

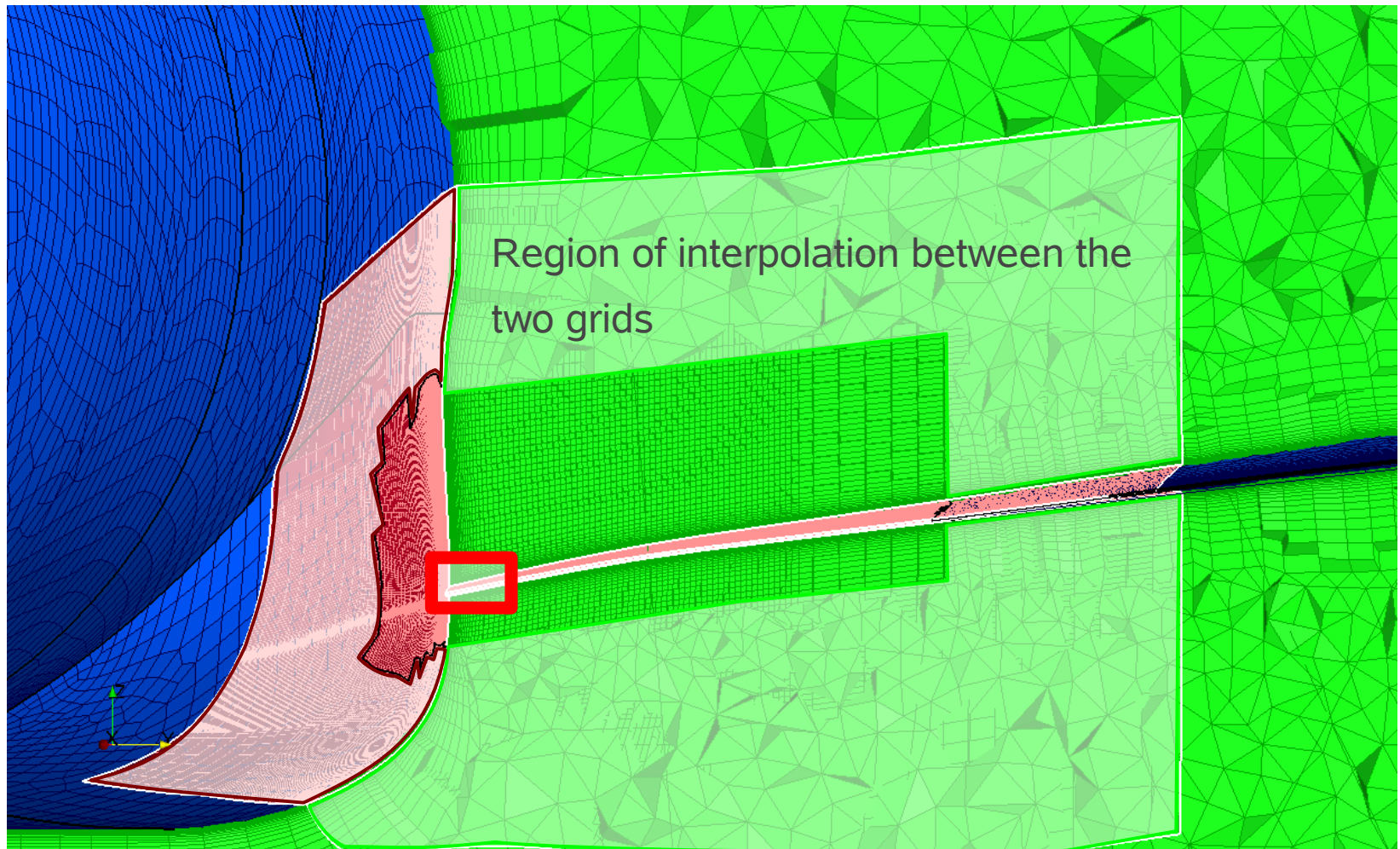




# Post-Workshop Activities

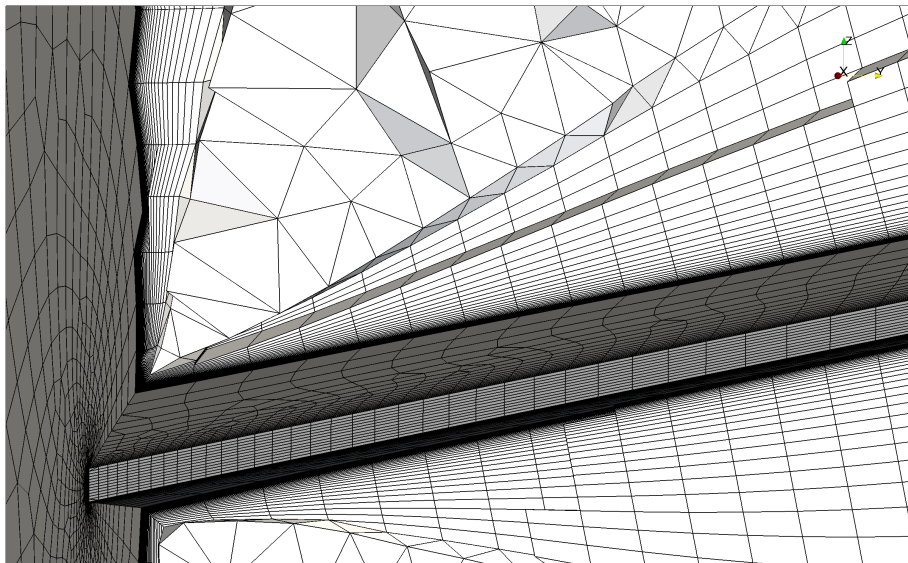


# Post-Workshop Activities

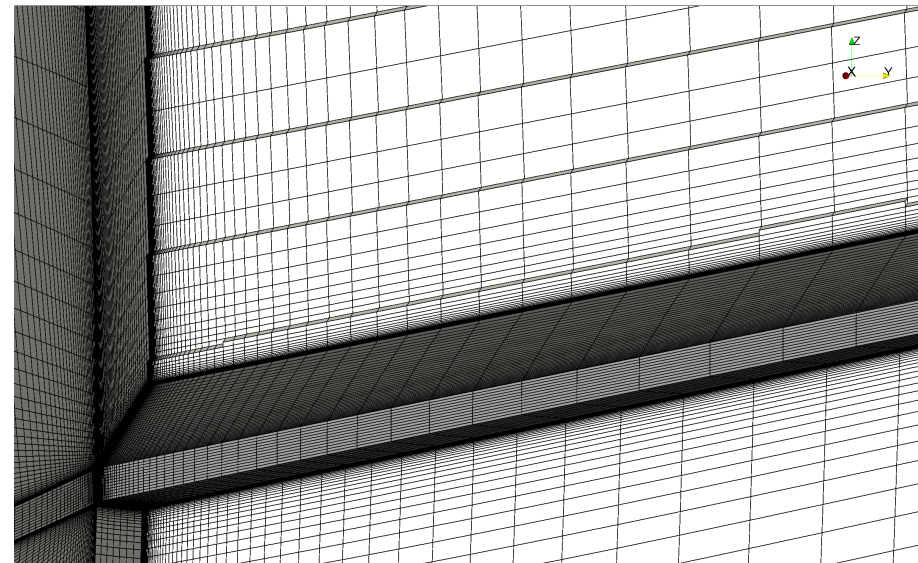


# Post-Workshop Activities

- Comparison between original Solar grid and hexa block of chimera grid, reveals the improved field discretization directly above the near-field layers



Original Solar grid

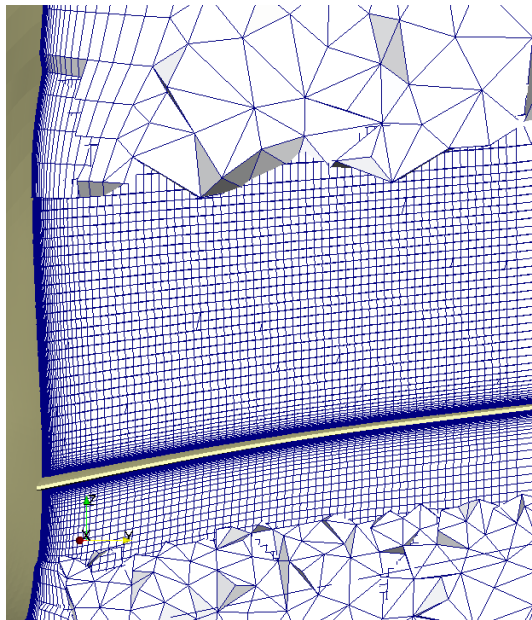


Hexa block of chimera grid

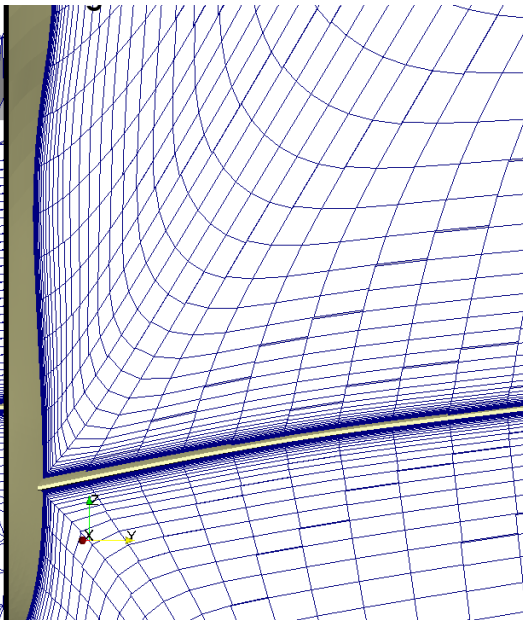


# Post-Workshop Activities

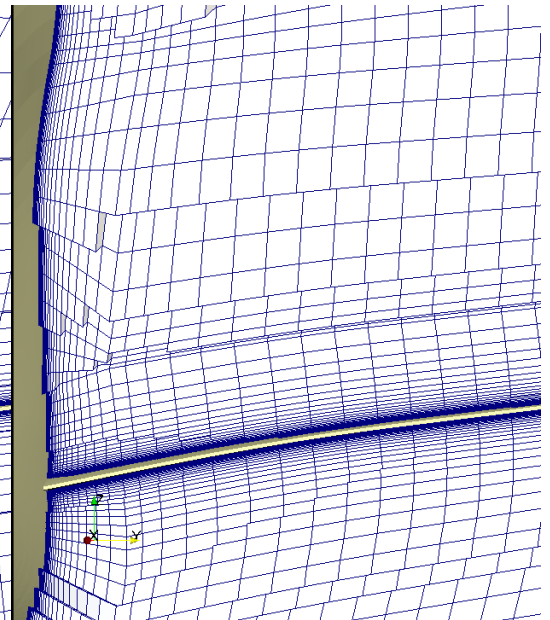
- Solar & hexa chimera grid is relatively fine, when compared to other (medium) structured grids



Solar & hexa block grid



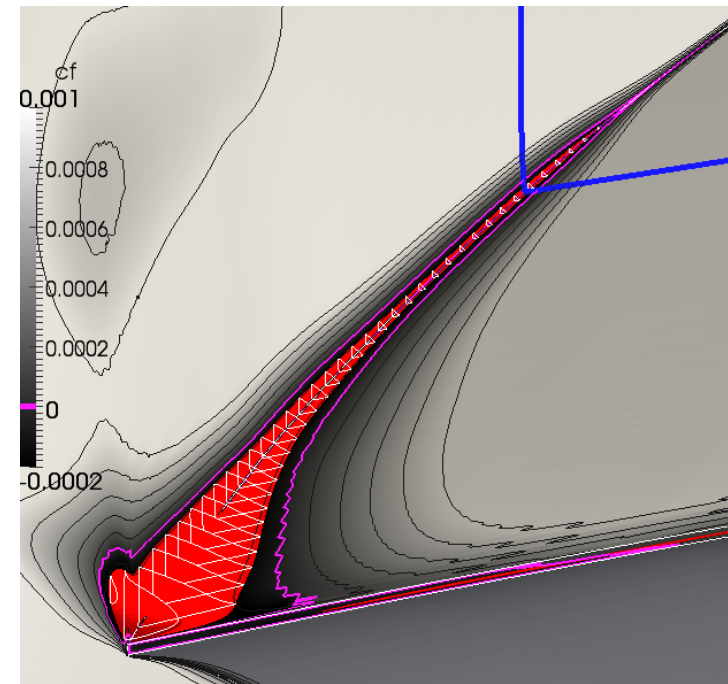
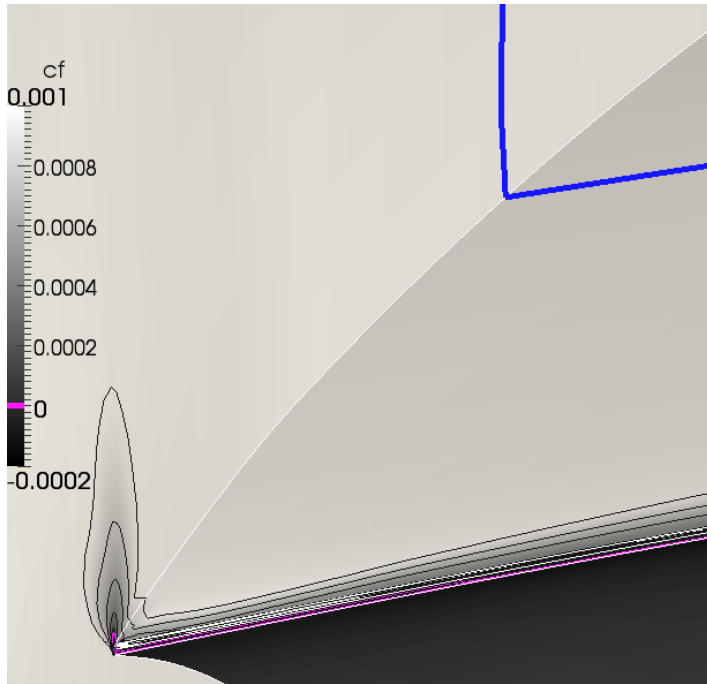
Boeing



ANSYS

# Post-Workshop Activities

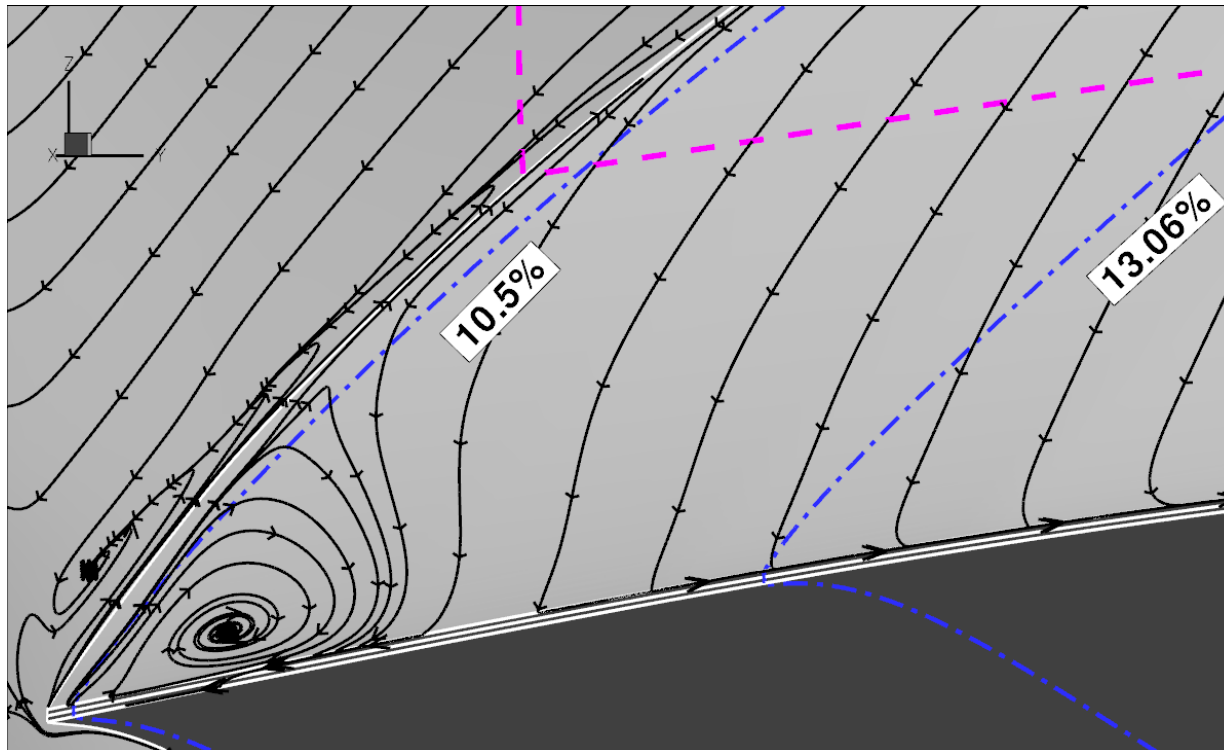
- Through improved field discretization, the separation bubble is resolved for target-C-lift = 0.5 computations (case 1.1;  $\alpha \approx 2.3^\circ$ )



- Solution on original Solar grid (left) reveals the same location of the shock (isocurve of critical pressure coefficient in blue) as on the chimera grid (right)
- Separation bubble (isosurface of x-velocity = -10 m/s in red) starts ahead of sonic line

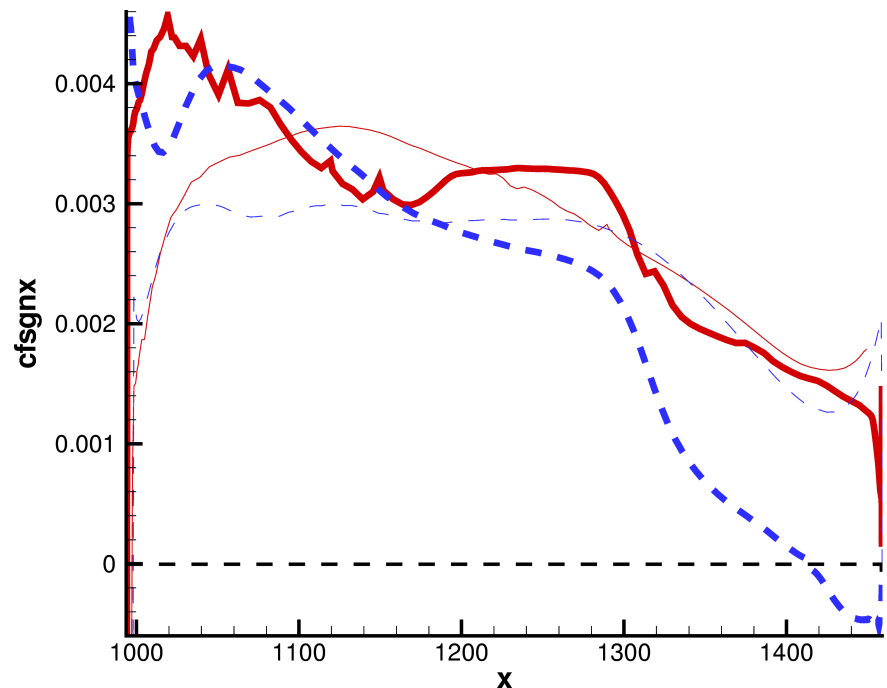
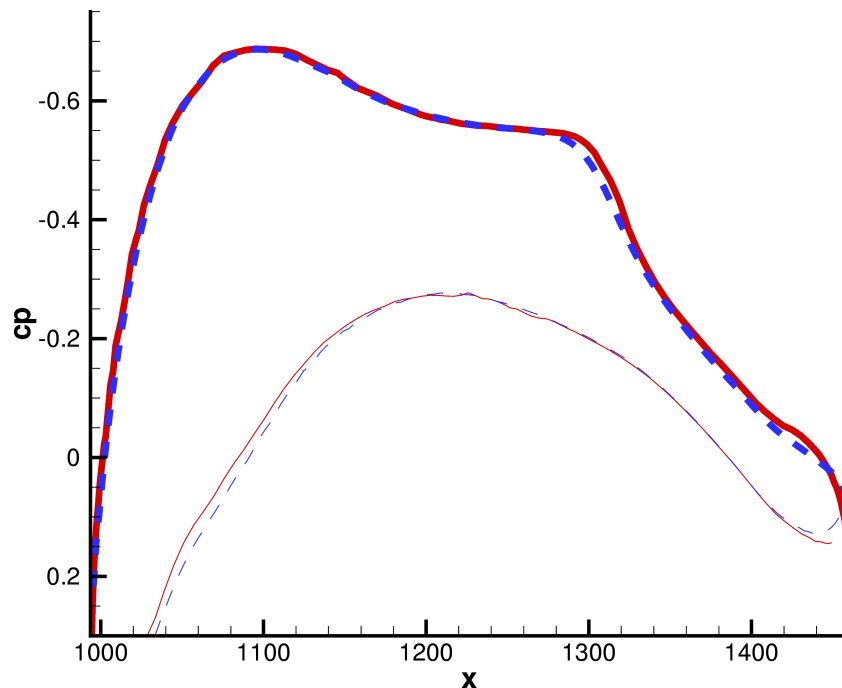
# Post-Workshop Activities

- Surface pressure and skin friction data compared at the innermost cut-plane of  $y/2b=10.5\%$
- First experimental cut-plane at  $y/2b=13.06\%$  outside of influence region of separation bubble



# Post-Workshop Activities

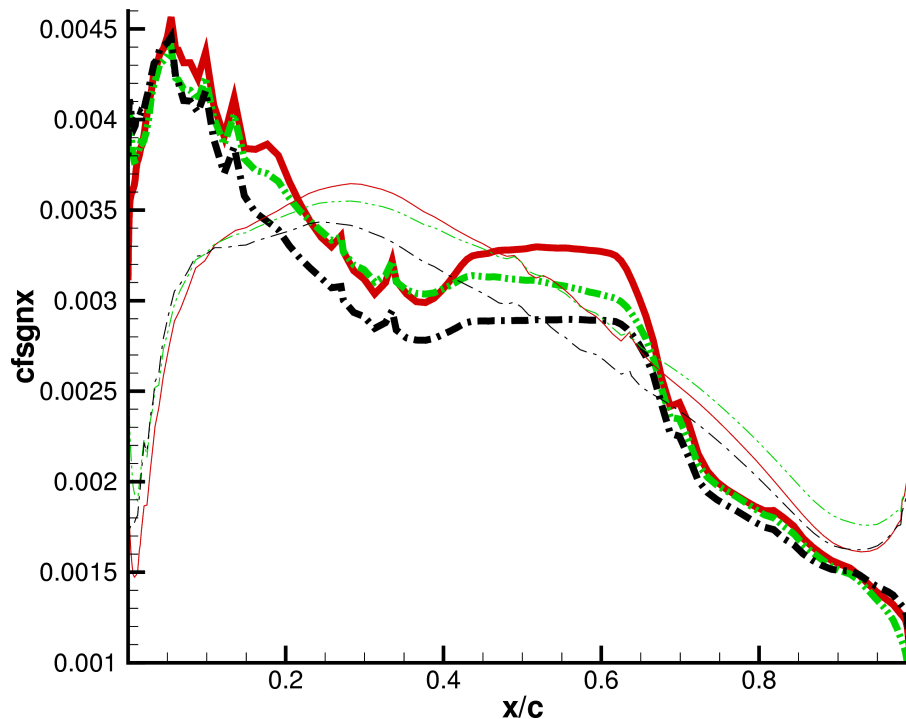
- No difference in pressure rise over shock between the two solutions
- Separation visible (only) through skin friction coefficient



Case 1.1; Solar grid (red, continuous) compared to chimera grid solution (blue, dashed) at  $y/2b=10.5\%$ ; pressure coefficient (left) and skin friction coefficient (right) for SA computations; upper wing data bold

# Post-Workshop Activities

- No separation found with other turbulence models on Solar grid (SA, SST, RSM)

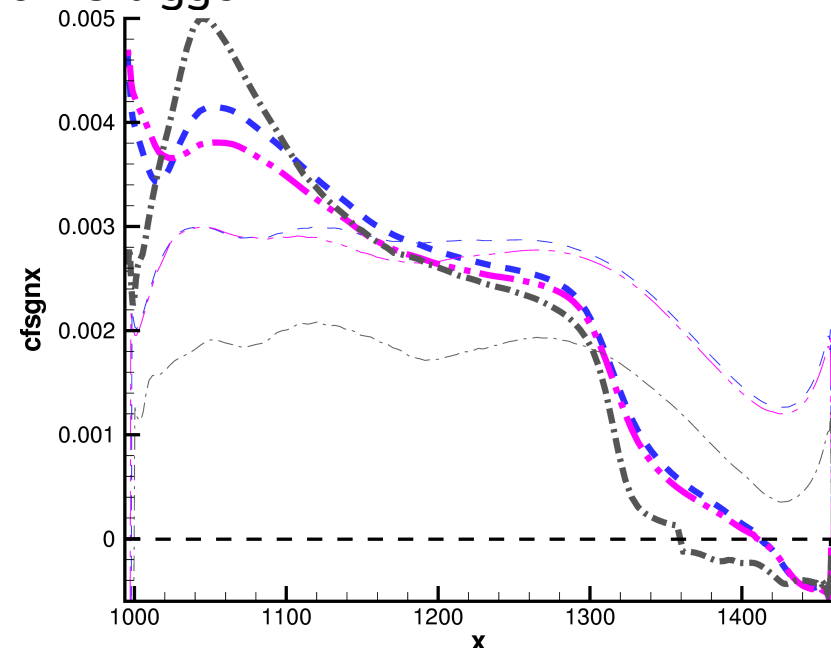
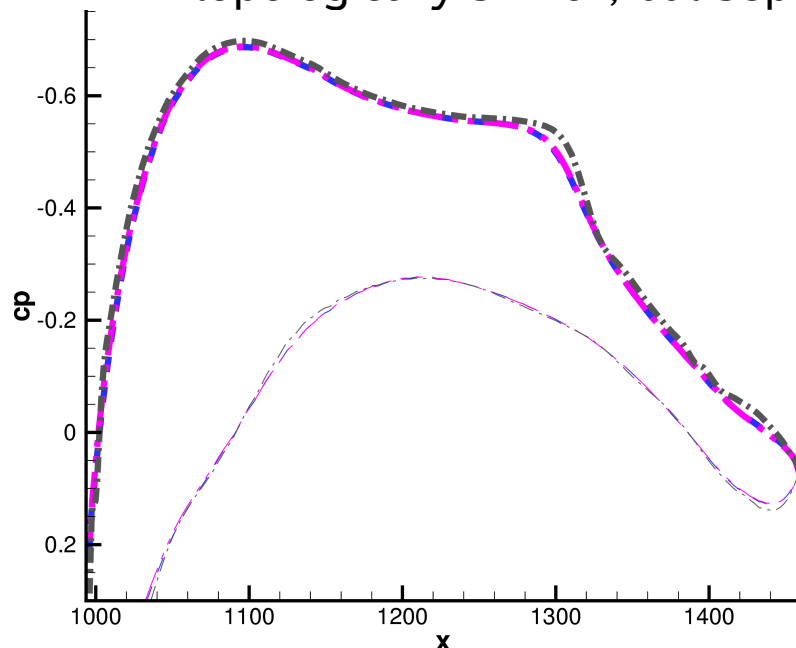


Case 1.1; Solar grid, SA (red, continuous), SST (green, dash-dot-dot) and RSM (black dash-dot) results at  $y/2b=10.5\%$ ; skin friction coefficient; upper wing data bold



# Post-Workshop Activities

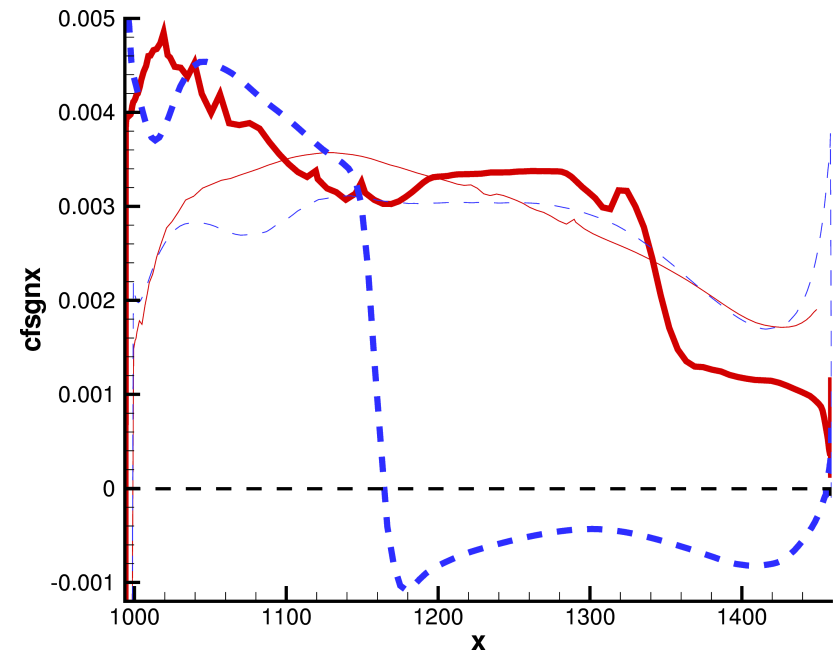
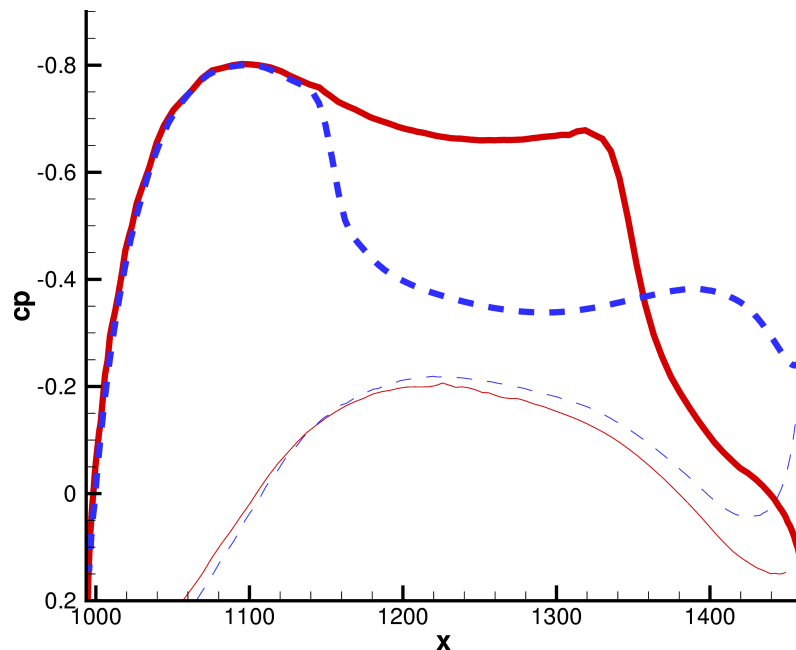
- On chimera grid, separation is found for all tested turbulence models (SA, SARC, RSM)
- No influence on separation of rotational corrections for SA; RSM topologically similar, but separation is bigger



Case 1.1; Chimera grid solution for SA (blue, dashed), SARC (pink, dash-dot-dot) and RSM (gray, dash-dot) at  $y/2b=10.5\%$ ; pressure coefficient (left) and skin friction coefficient (right); upper wing data bold

# Post-Workshop Activities

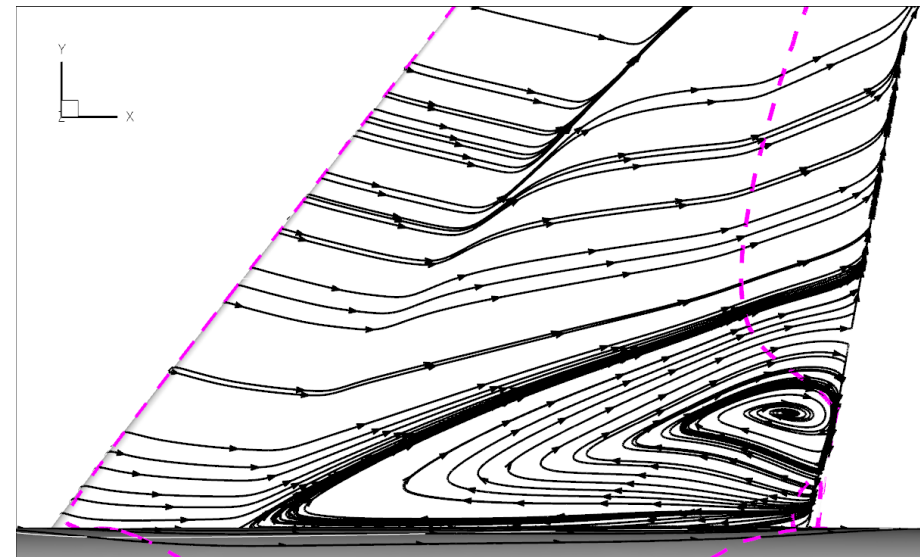
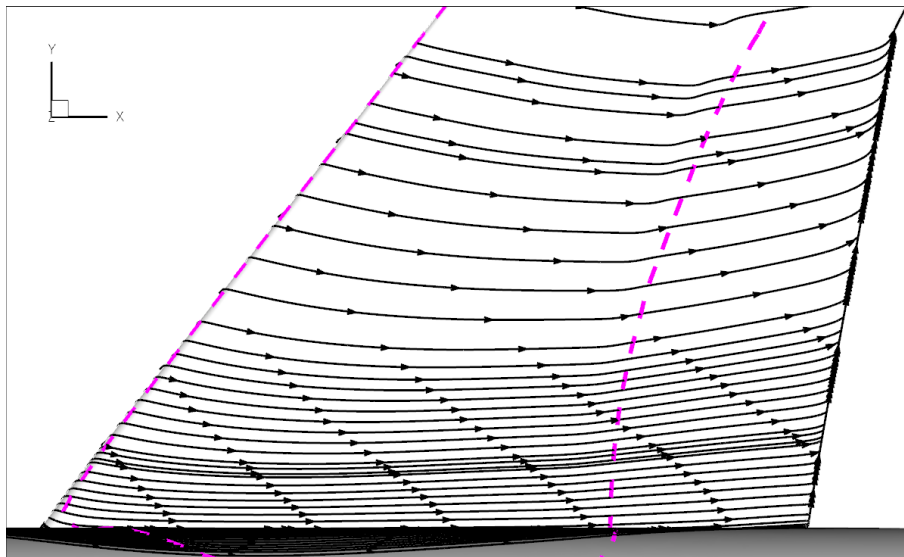
- At the edge of the envelope ( $\text{aoa} = 4^\circ$ ), the inferior junction discretization of the standard Solar grid has (more) severe effects
- Begin of separation moves upstream



Case 1.2,  $\text{aoa} = 4^\circ$ ; Solar grid (red, continuous) compared to chimera grid solution (blue, dashed) at  $y/2b=10.5\%$ ; pressure coefficient (left) and skin friction coefficient (right); upper wing data bold

# Post-Workshop Activities

- Massive separation bubble dominates inner wing flowfield in chimera grid solution



Case 1.2,  $\text{aoa} = 4^\circ$ ; Solar grid (left) compared to chimera grid solution (right); critical pressure coefficient (pink, dashed curve) and skin friction lines

# Conclusions

- Grid family generation process developed for Solar, applicable to any advancing-layer/front method
  - Satisfactory grid convergence; behavior (second order) and gradient
  - Even without resolution of separation bubble at SOB (not resolved on any grid level)
- Performed successful adjoint-based grid assessment and improvement
- Grid deficiency for junction flow discretization solved via chimera grid technique
  - At  $Cl=0.5$ , separation onset seemingly triggered by shock destabilizing effect on junction boundary layer
  - At  $aoa=4^\circ$ , substantial change in wing root flow topology